

GRAINE計画：

気球搭載型大口径エマルジョン望遠鏡による
10MeV-100GeVガンマ線の高解像度・偏光観測

高橋 覚 (神戸大)

for GRAINE collaboration

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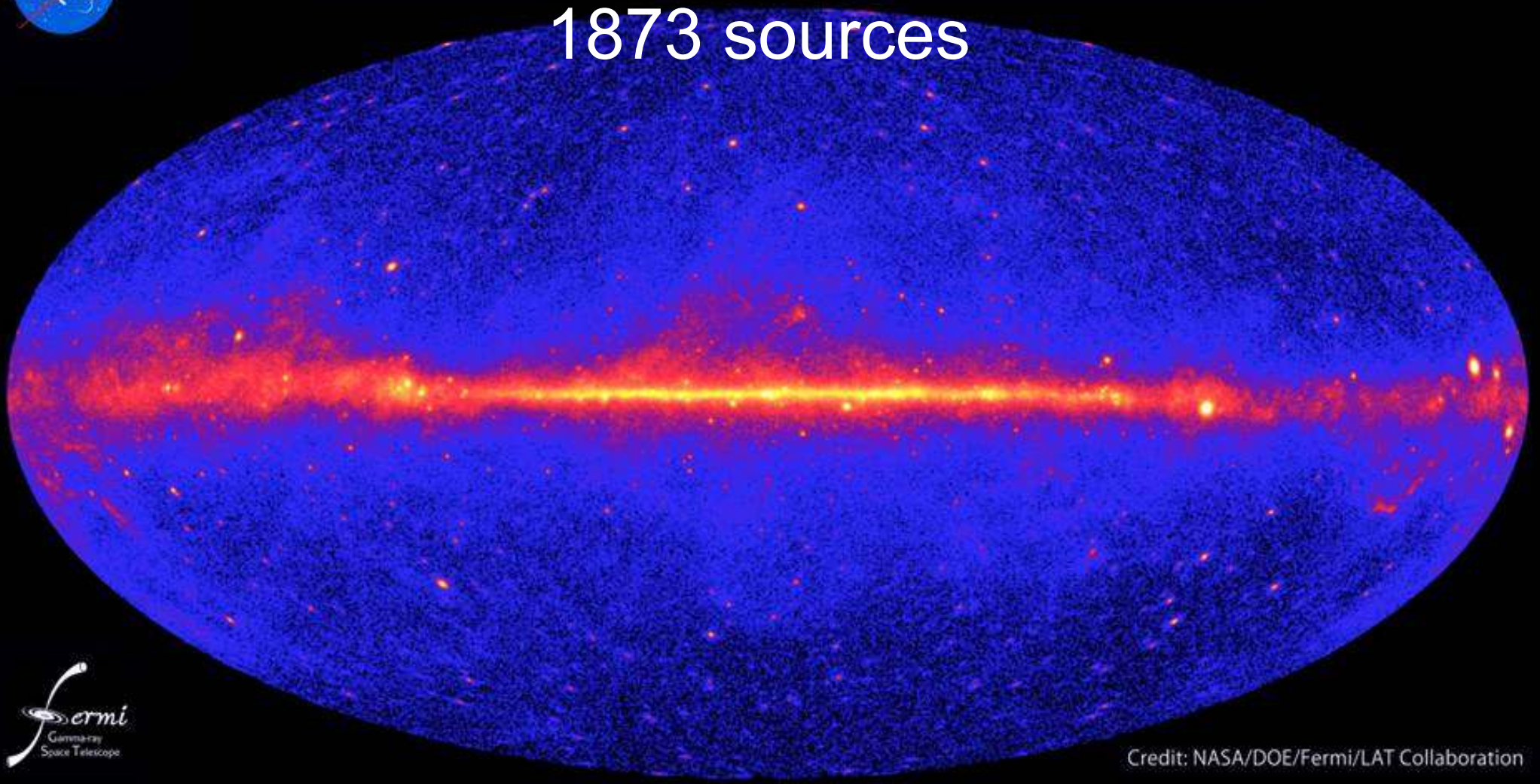
(1)Kobe University, (2)ISAS/JAXA, (3)Nagoya University, (4)Okayama University of science,
(5)Aichi University of education, (6)Utsunomiya University

PI : 青木 茂樹 (神戸大)



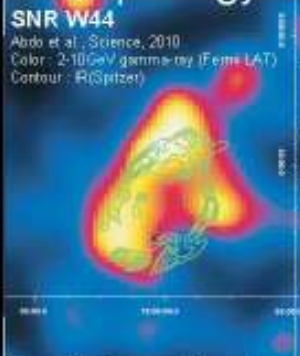
Fermi two-year all-sky map ($E_\gamma > 1\text{GeV}$)

1873 sources

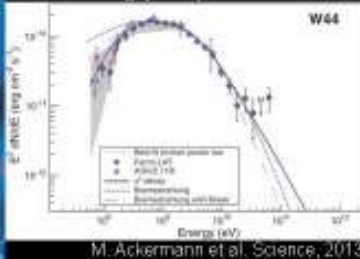


Credit: NASA/DOE/Fermi/LAT Collaboration

Morphology



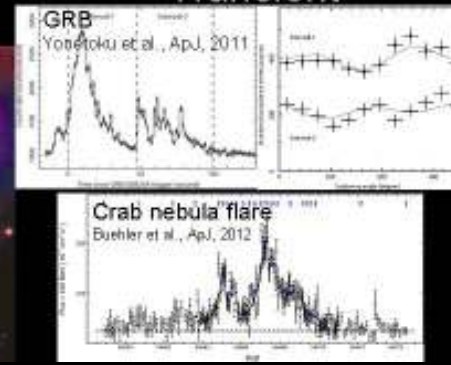
Energy spectrum



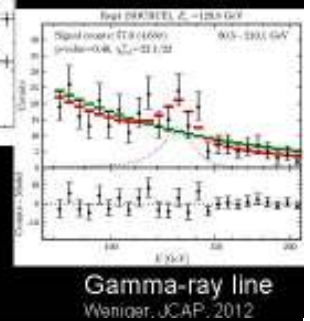
Polarization



Transient

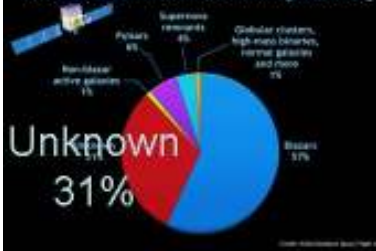


Dark matter



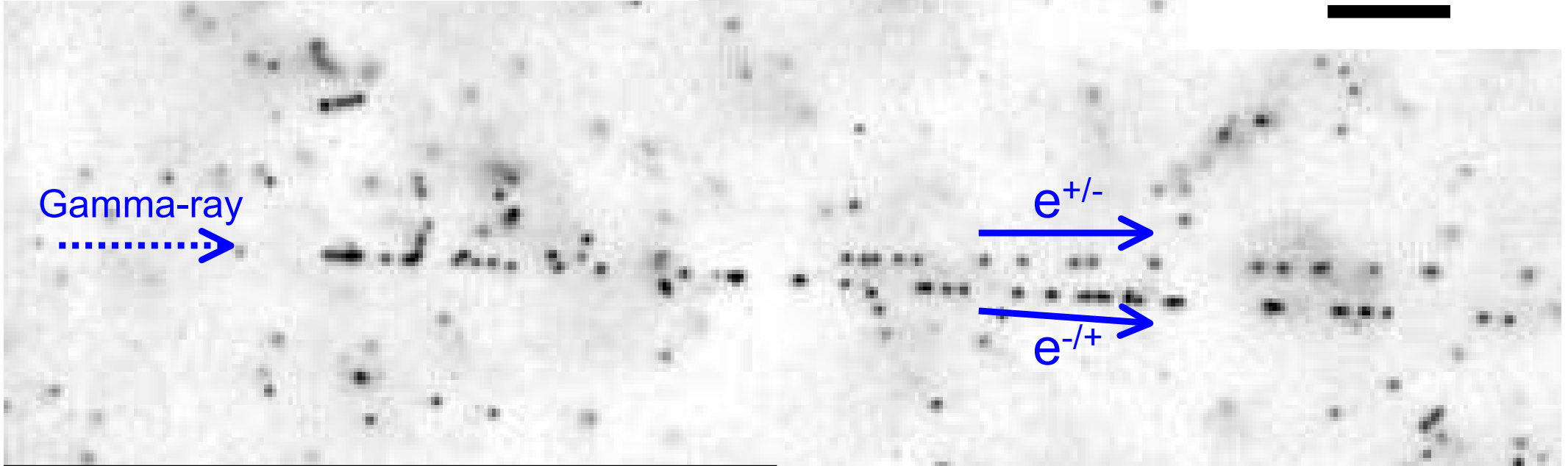
Un-ID

What has Fermi found: The LAT two-year catalog

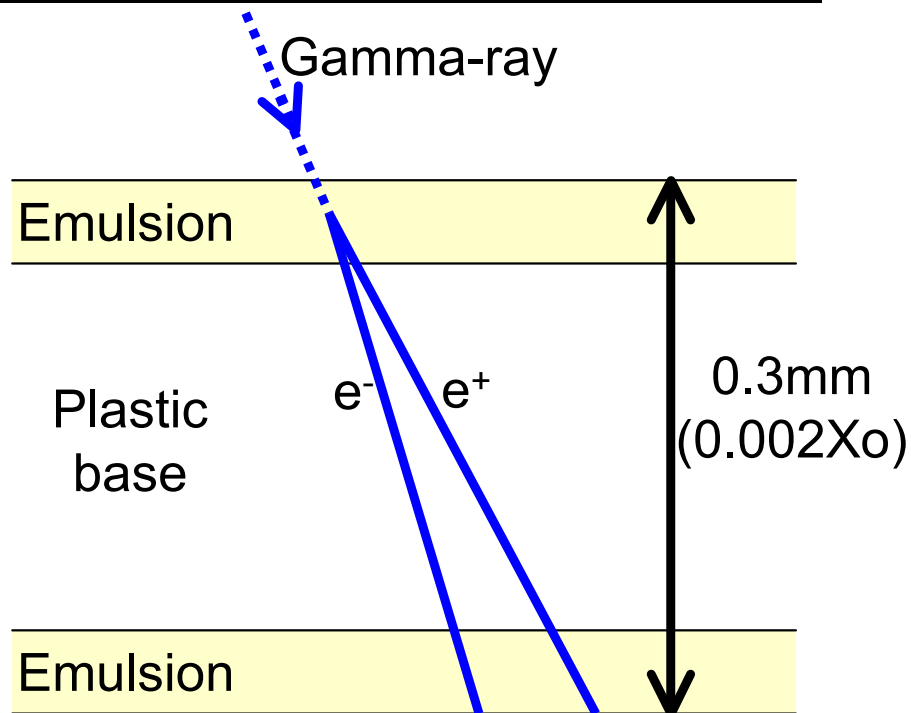


Nuclear emulsion

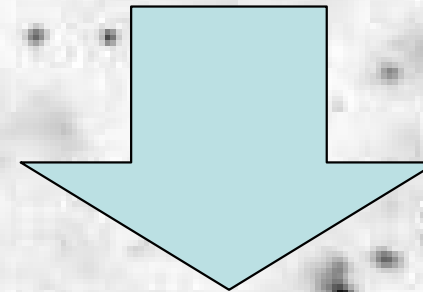
Microscopic view
10micron



Cross sectional view of an emulsion film



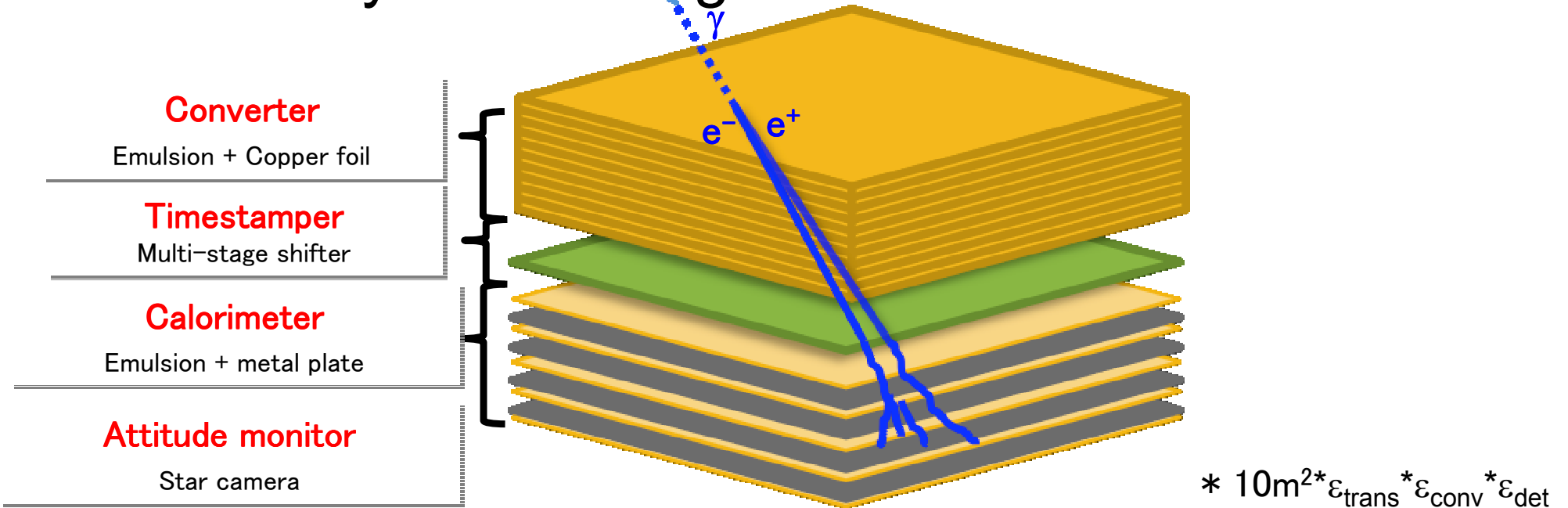
Powerful tracking device
>High spatial resolution : ~ 1 micron
>Small radiation length : $0.002X_0$



High angular resolution for gamma-ray
Sensitive to gamma-ray polarization

GRAINE

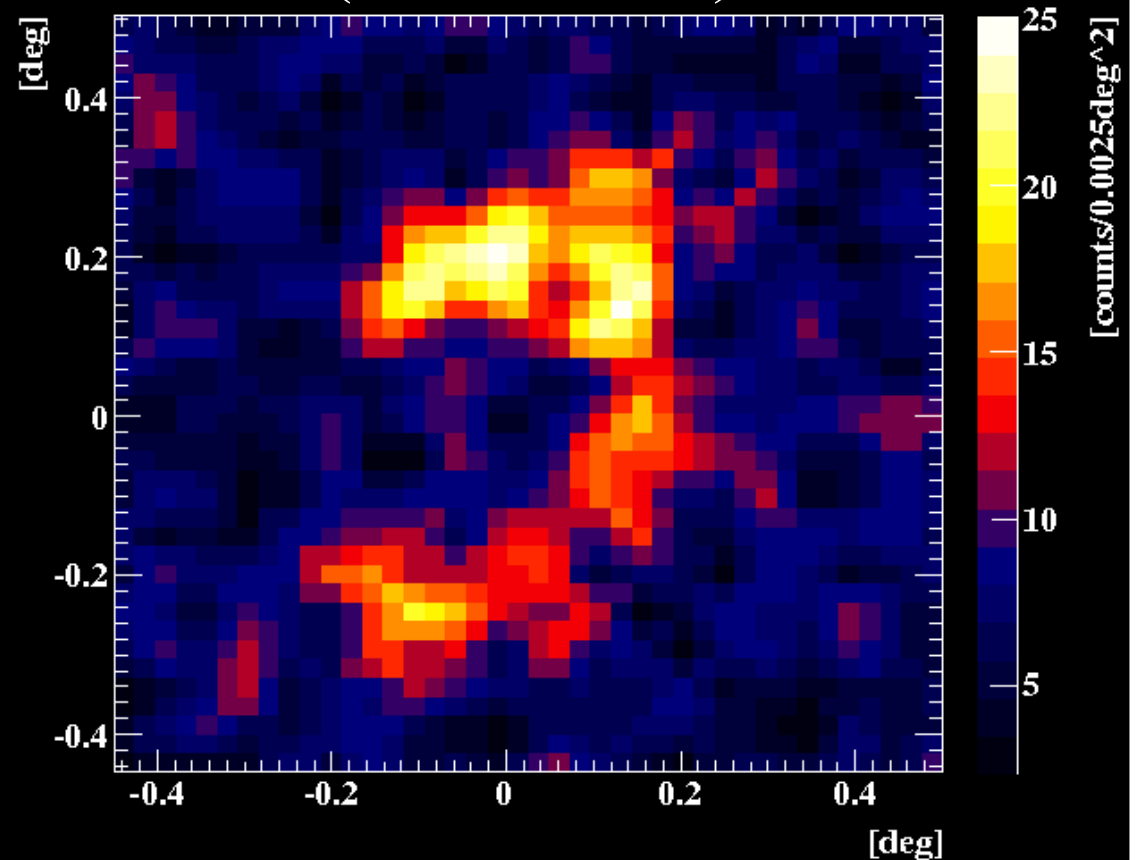
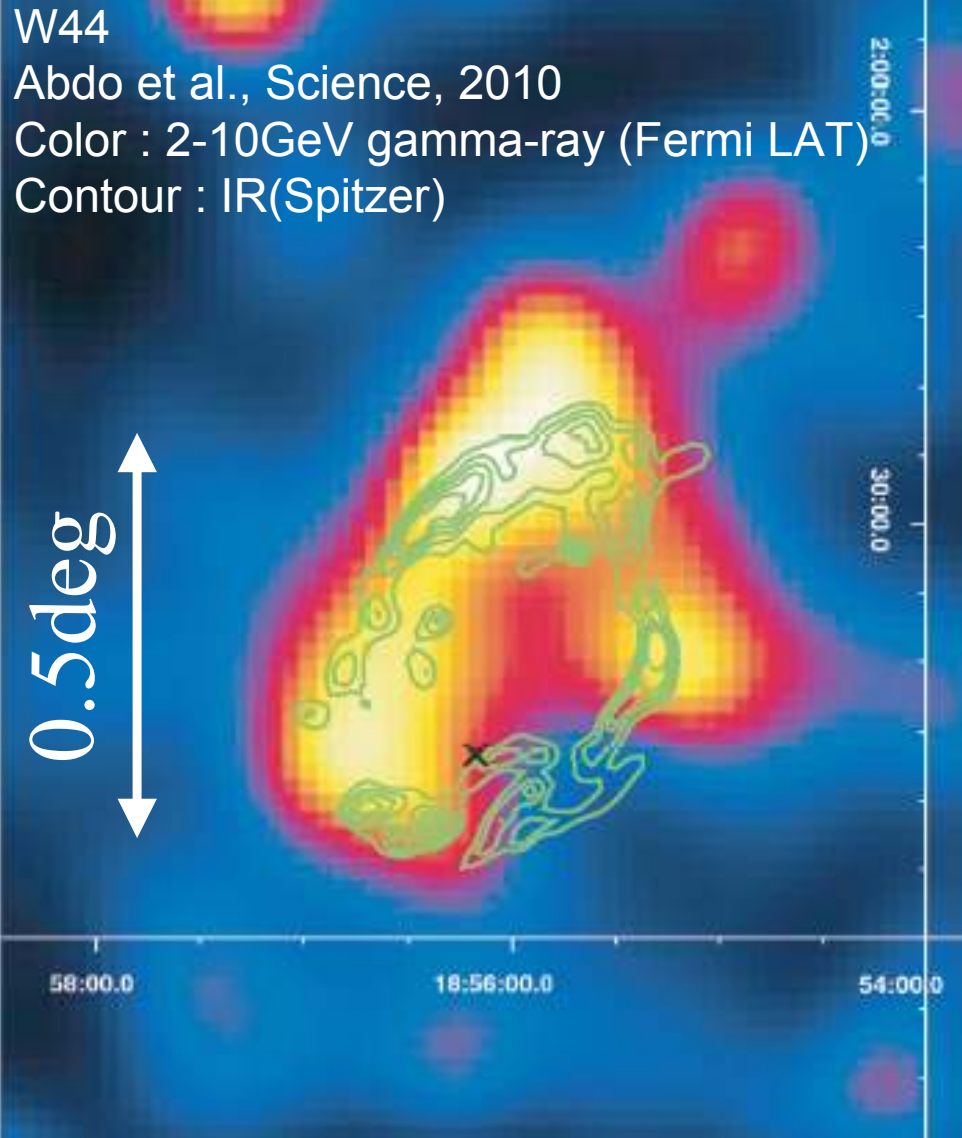
Gamma-Ray Astro-Imager with Nuclear Emulsion



	Fermi LAT	GRAINE
Angular resolution @100MeV	6.0deg (105mrad)	1.0deg (17mrad)
@1GeV	0.90deg (16mrad)	0.1deg (1.7mrad)
Energy range	20MeV – 300GeV	10MeV – 100GeV
Polarization sensitivity	No	Yes
Effective area @ 100MeV	0.25m ²	2.1m ² *
@ 1GeV	0.88m ²	2.8m ² *
Dead time	26.5 μ sec _(readout time)	Dead time free

High resolution imaging

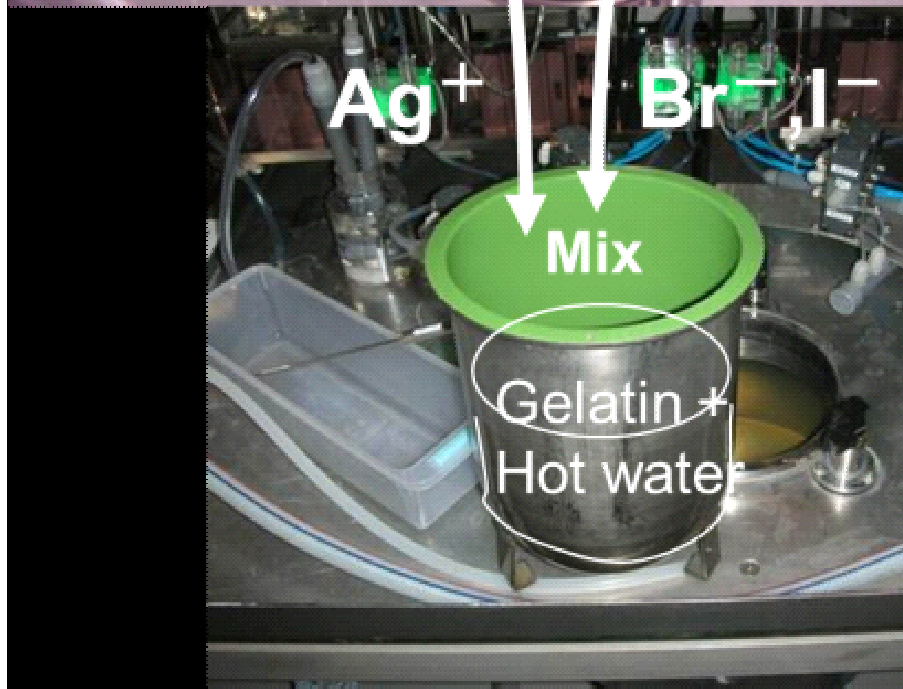
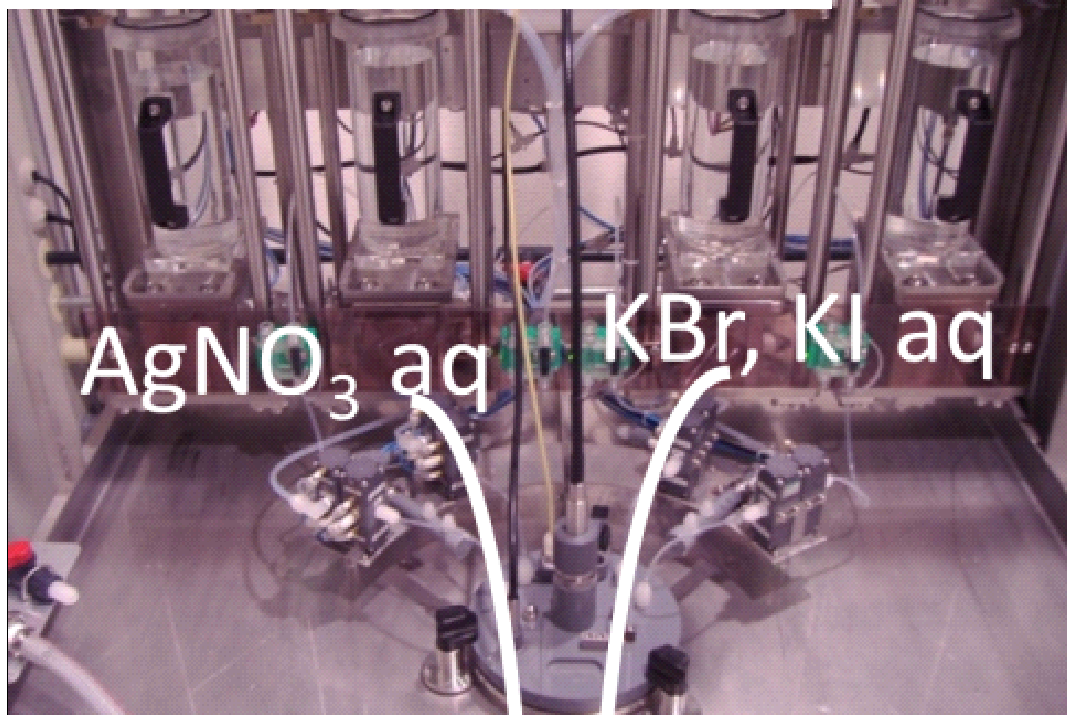
GRAINE (Simulation)



- **3flights** (41.7m²days)
- >1GeV
- Smearing IR(Spitzer) distribution
with 0.08deg(1.4mrad)
- Considering atmospheric gamma-ray(>1GeV) as BG

Emulsion production

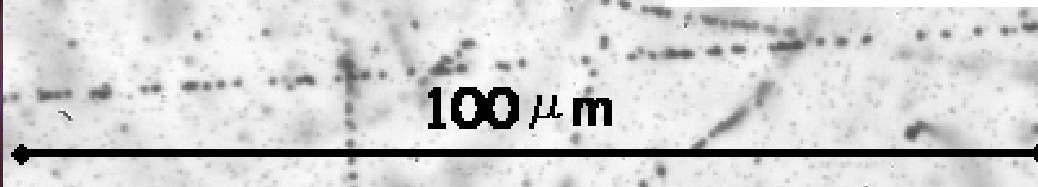
Fuji Janet Co., Ltd., Nagoya Univ.



Gelatin 1/4, Na type, Fe x 2, MIP (XAA, 20deg, 40min)

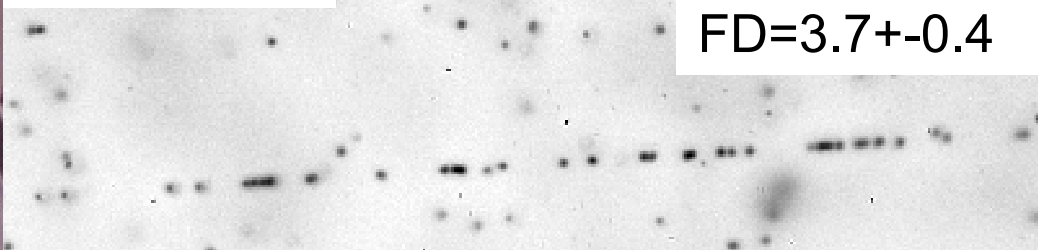
New type

GD=86.1±4.7
FD=2.9±0.9

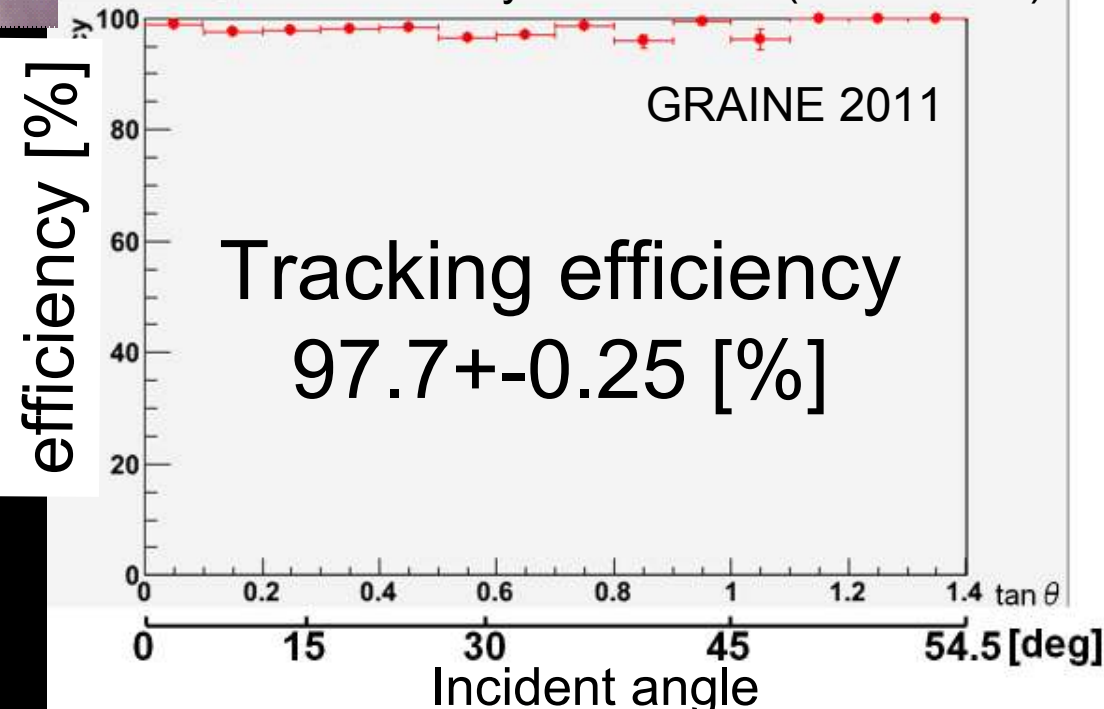


OPERA type

GD=34.8±0.6
FD=3.7±0.4



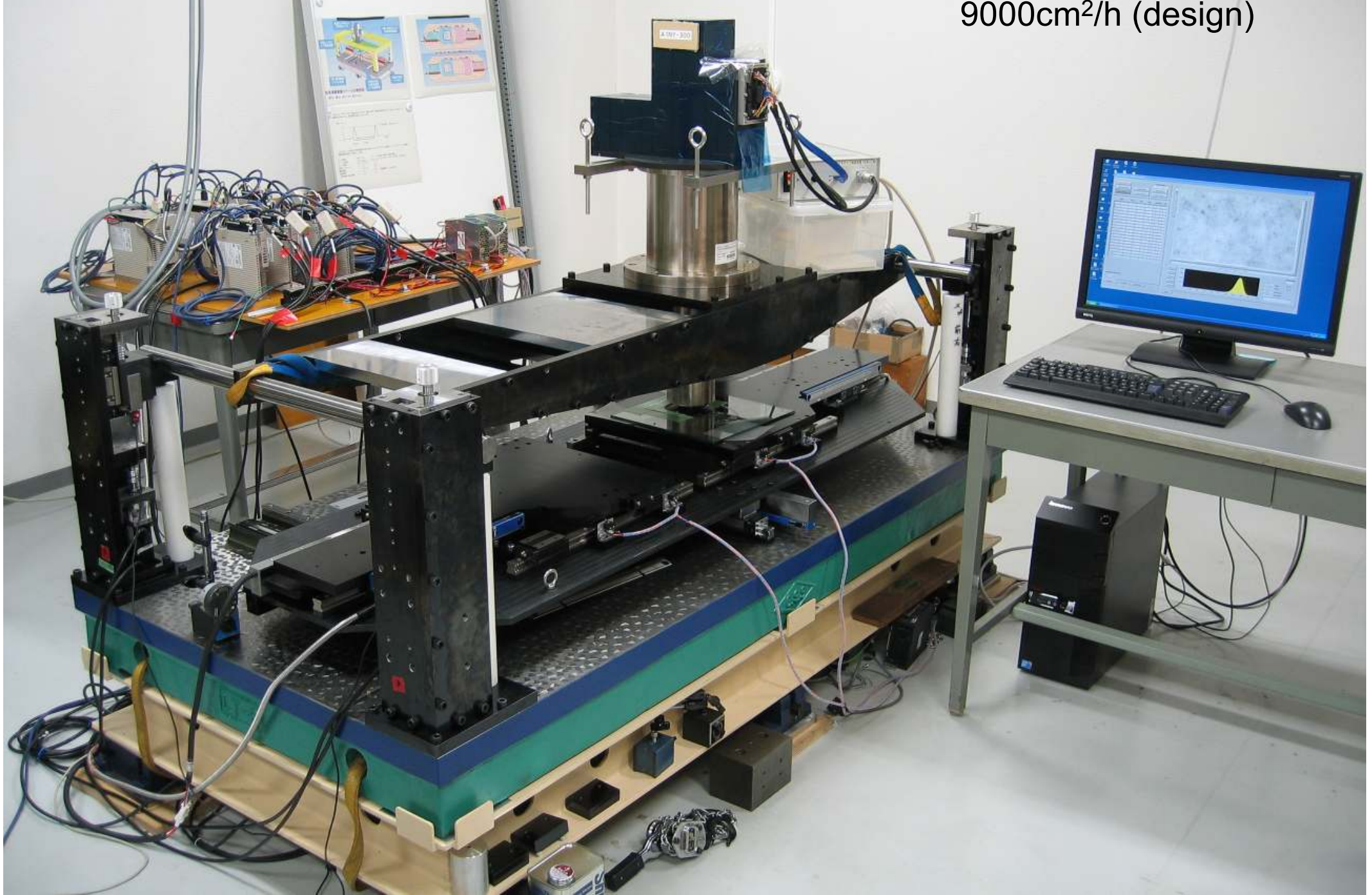
efficiency Evaluated by K. Kamada (Kobe Univ.)



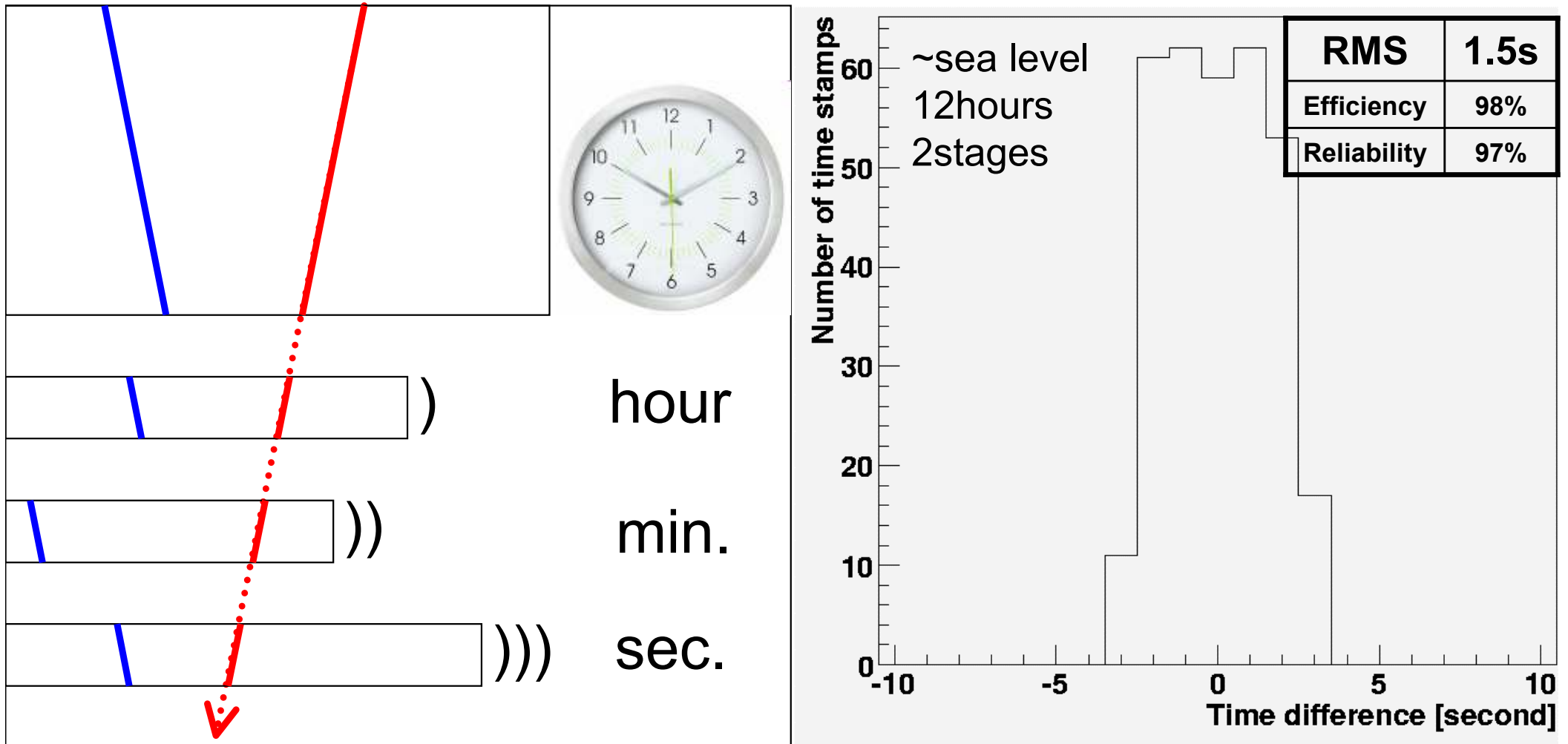
Automated emulsion read-out system (Nagoya Univ.)

Hyper-TS: Next Generation Read-out system

Scanning speed
9000cm²/h (design)



Multi-stage shifter



S.Takahashi et al., Nucl. Instr. And Meth. A, 620 (2010) 192-195

Consists of emulsion films with Small R.L., High pos & ang resolution

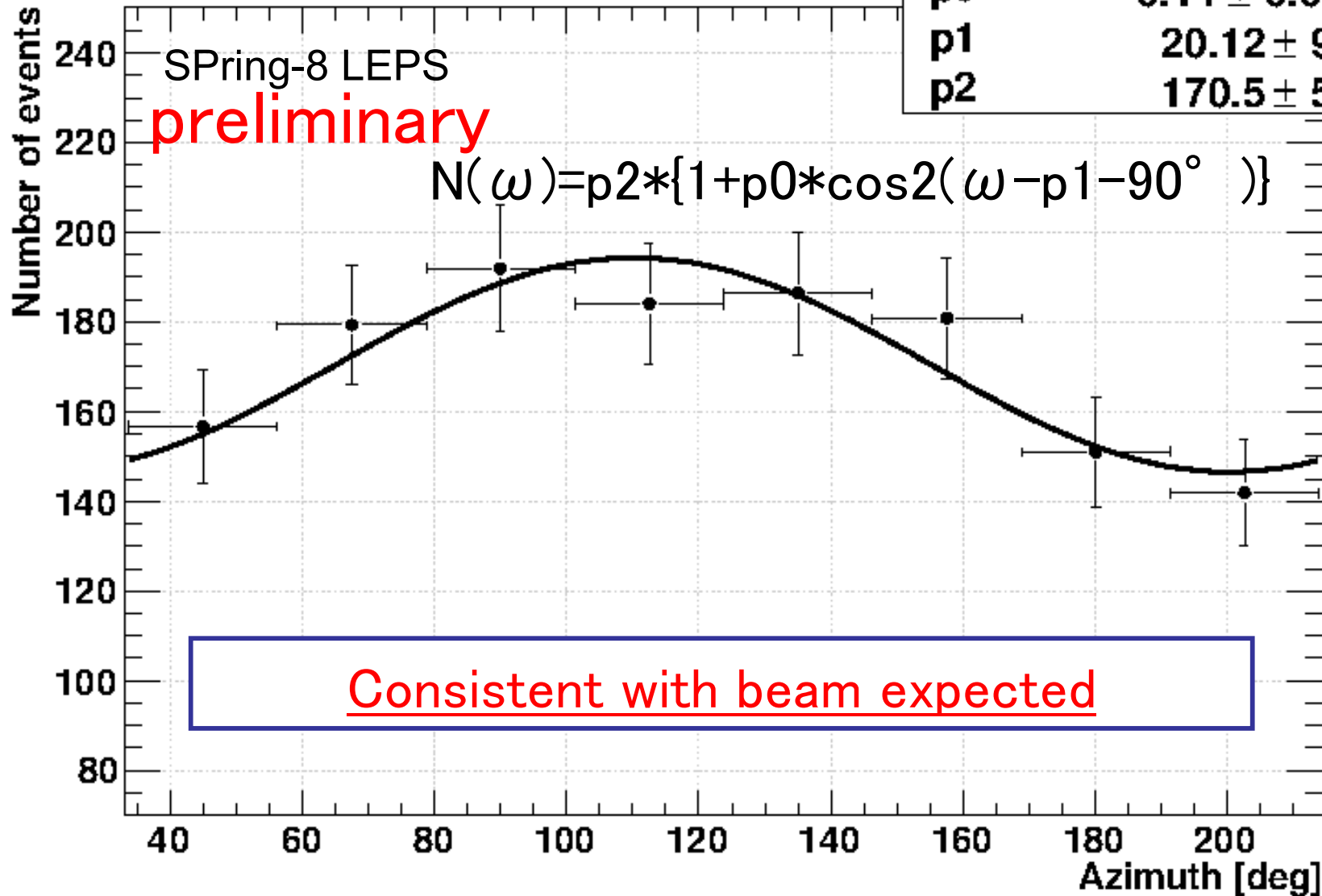
→ Low P threshold, High reliability, High efficiency, Large scale

Simple component, Compact, Light, HV free, Low power consumption, Dead time free

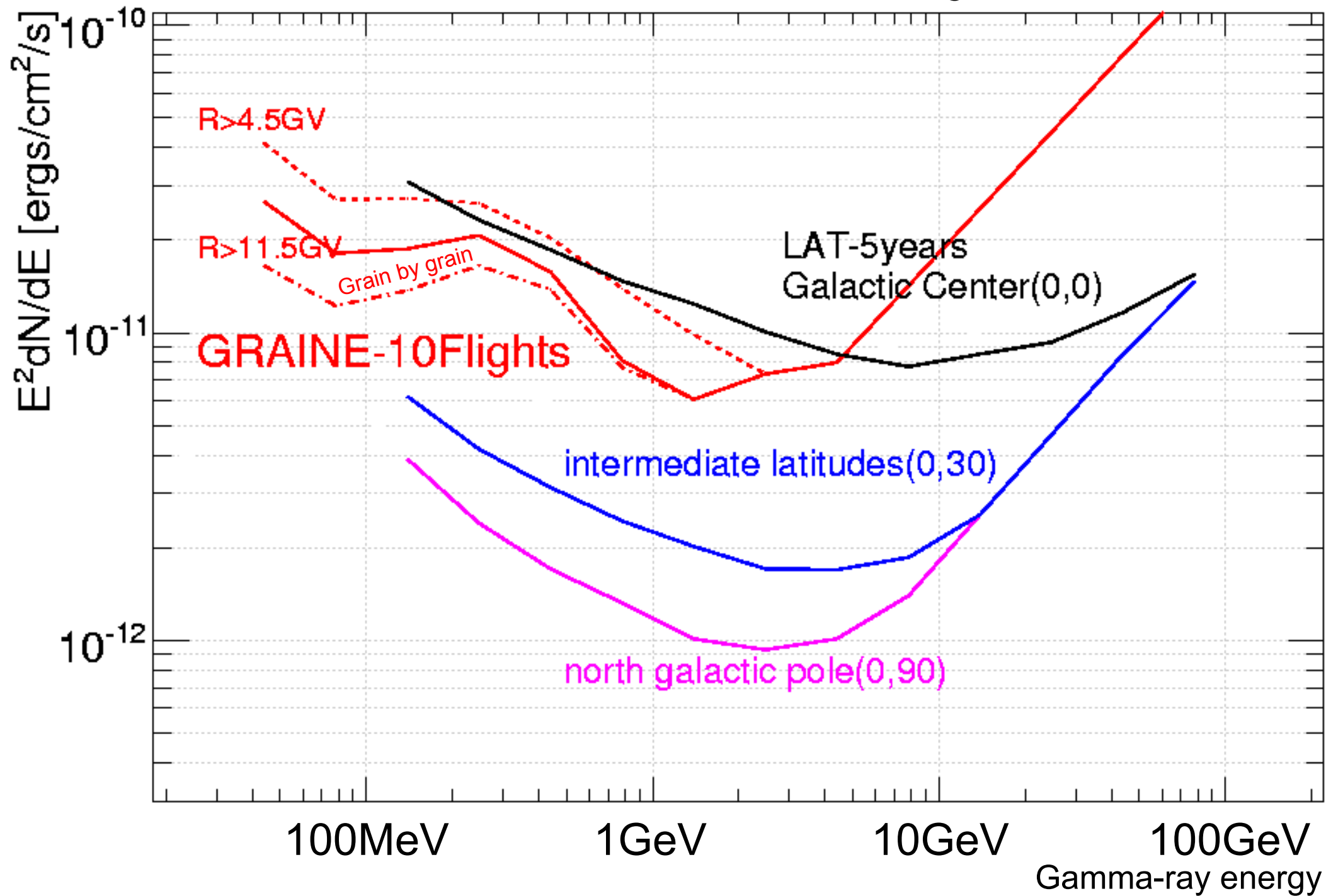
Polarization sensitivity

Azimuthal Distribution

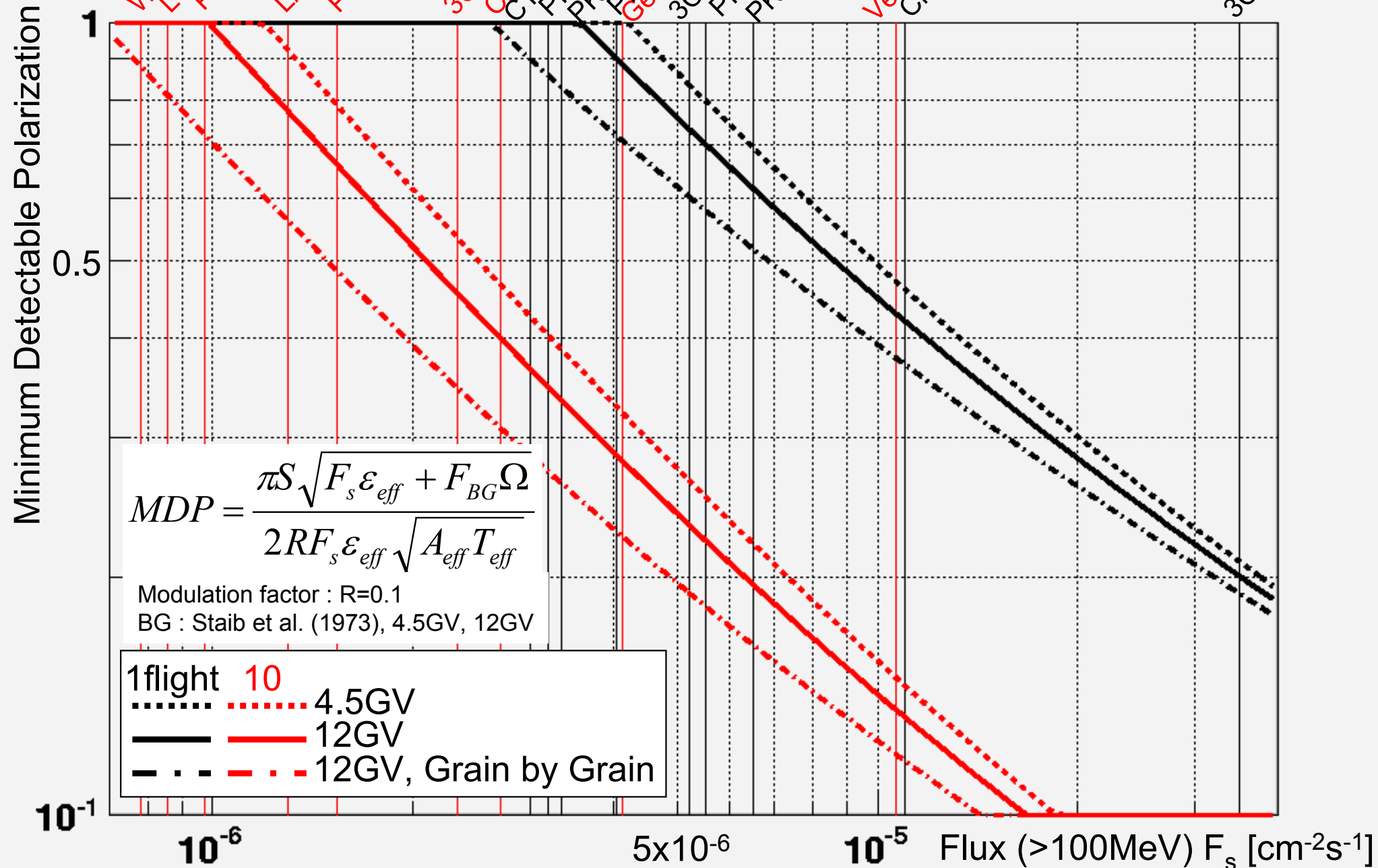
χ^2 / ndf	1.532 / 5
p0	0.14 ± 0.03962
p1	20.12 ± 9.506
p2	170.5 ± 5.089



Source sensitivity



Polarization sensitivity



$$MDP = \frac{\pi S \sqrt{F_s \epsilon_{eff} + F_{BG} \Omega}}{2 R F_s \epsilon_{eff} \sqrt{A_{eff} T_{eff}}}$$

Modulation factor : R=0.1
 BG : Staib et al. (1973), 4.5GV, 12GV

- 1flight 10
- 4.5GV
- 12GV
- . - 12GV, Grain by Grain

GRAINE roadmap

- 8th/June/2011, TARF, JAXA Scientific Ballooning, PI : S. Aoki (Kobe Univ.)
12.5cm x 10cm aperture area, 4.3hours (1.6hours@35km) flight duration
 - Working test for each component
 - Connection test between components
 - Measurement of atmospheric gamma-rays
- 2014(Planned), Alice Springs, JAXA International Scientific Ballooning
2500cm² aperture area, 1 day flight duration
 - Overall test by detecting known gamma-ray source
 - Observation with highest imaging resolution
- 2015–
10m² aperture area, 7days flight duration
 - Starting scientific observation

Emulsion chamber

Aperture area : 12.5cm x 10cm

◆ Flatness compensation films

OPERA film x 2

◆ Converter

102 emulsion films, 91 copper foils (50 μ m)

1.1kg, 35.0mm, 0.54Xo ($\epsilon_{\text{conv}} = 34\%$)

OPERA film x 10 (go-ban part)

OPERA film x 88 + Copper foil x 88

New type gel film x 4 + Copper foil x 3

-Target & Detector

-Precise measurement of incident direction

-0.08deg@1-2GeV, 0.93deg@100MeV

-Measurement of gamma-rays polarization

-Interface of timestamper

-Energy measurement of gamma-rays (<~GeV)

-Momentum measurement of electron pair with MCS

◆ Timestamper

1st : OPERA film x 2

2nd : OPERA film x 2

3rd : OPERA film x 2

New type gel film x 1

◆ Calorimeter

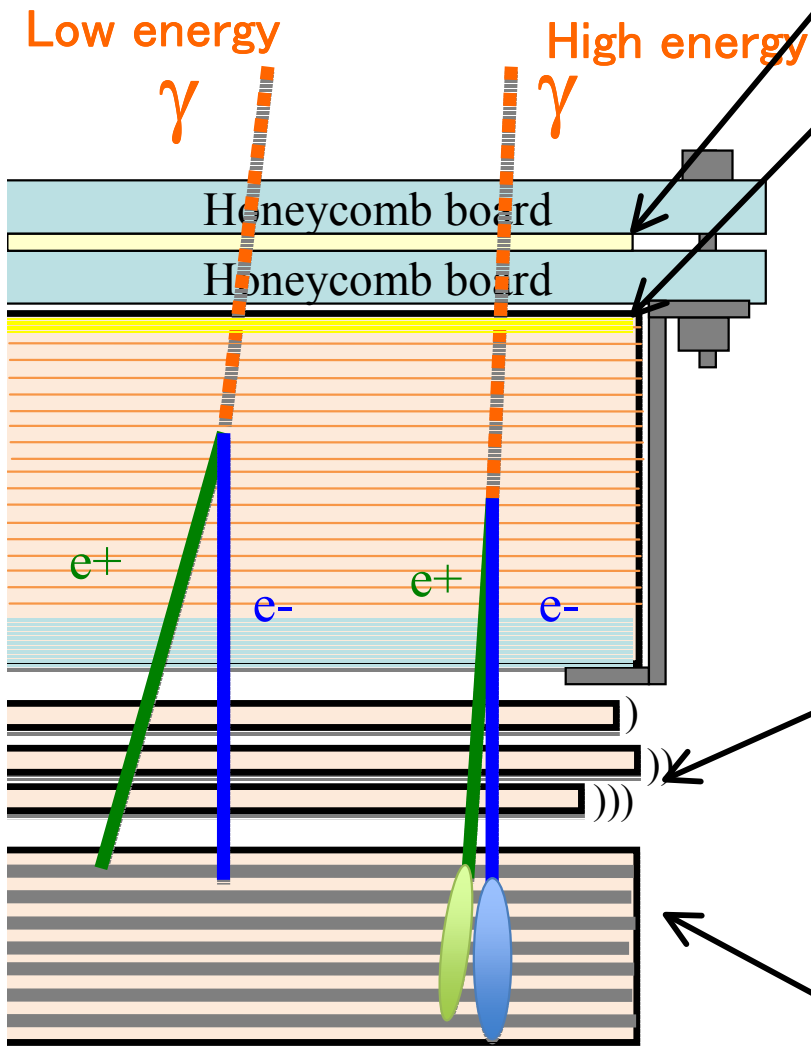
32 emulsion films, 10 (0.5mm) & 17 (1mm) lead plates

3.2kg, 31.9mm, 4.0Xo (Shower Max.@ a few GeV)

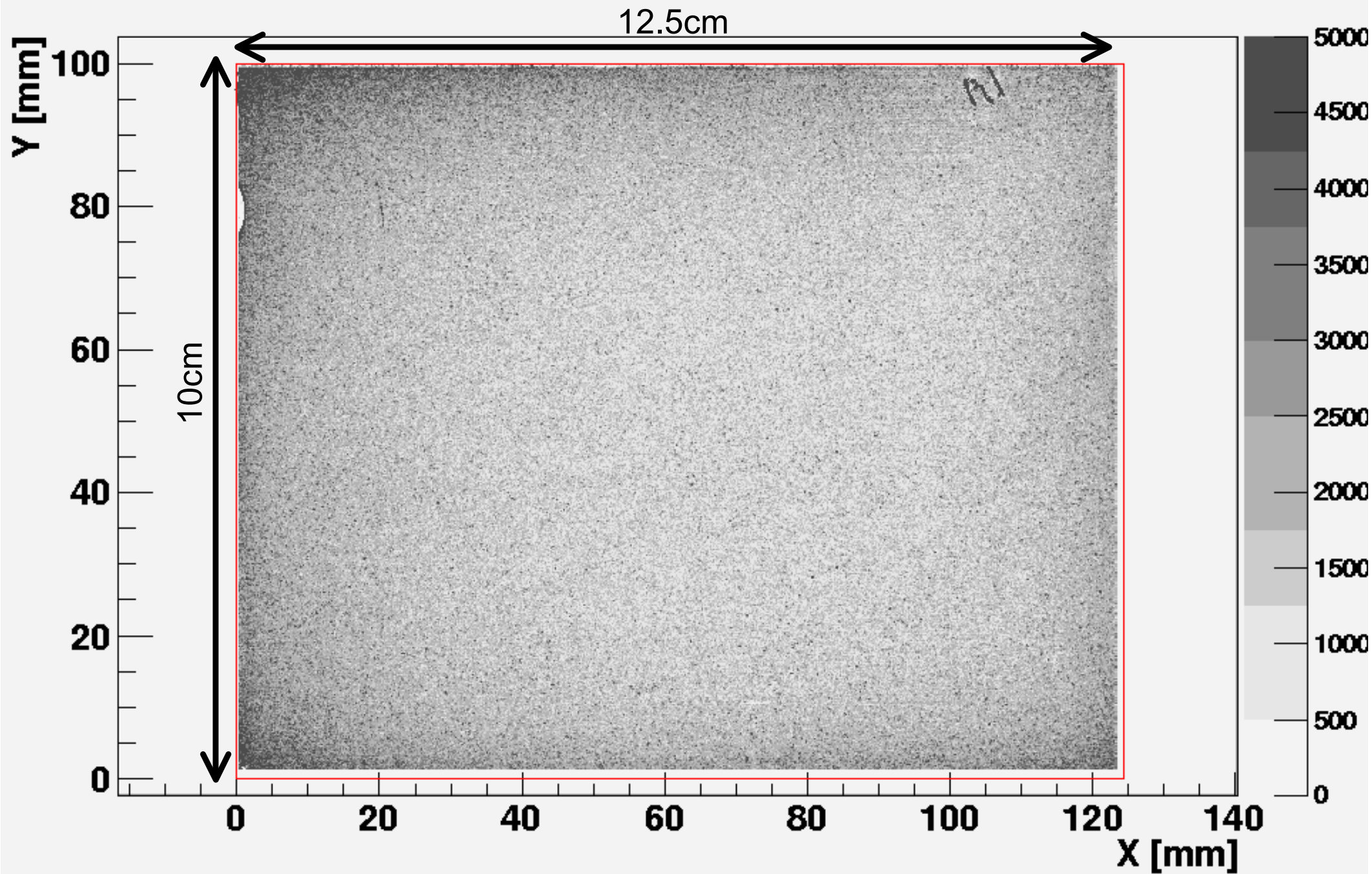
OPERA film x 5 (go-ban part)

(OPERA film + lead plate(0.5mm))x10

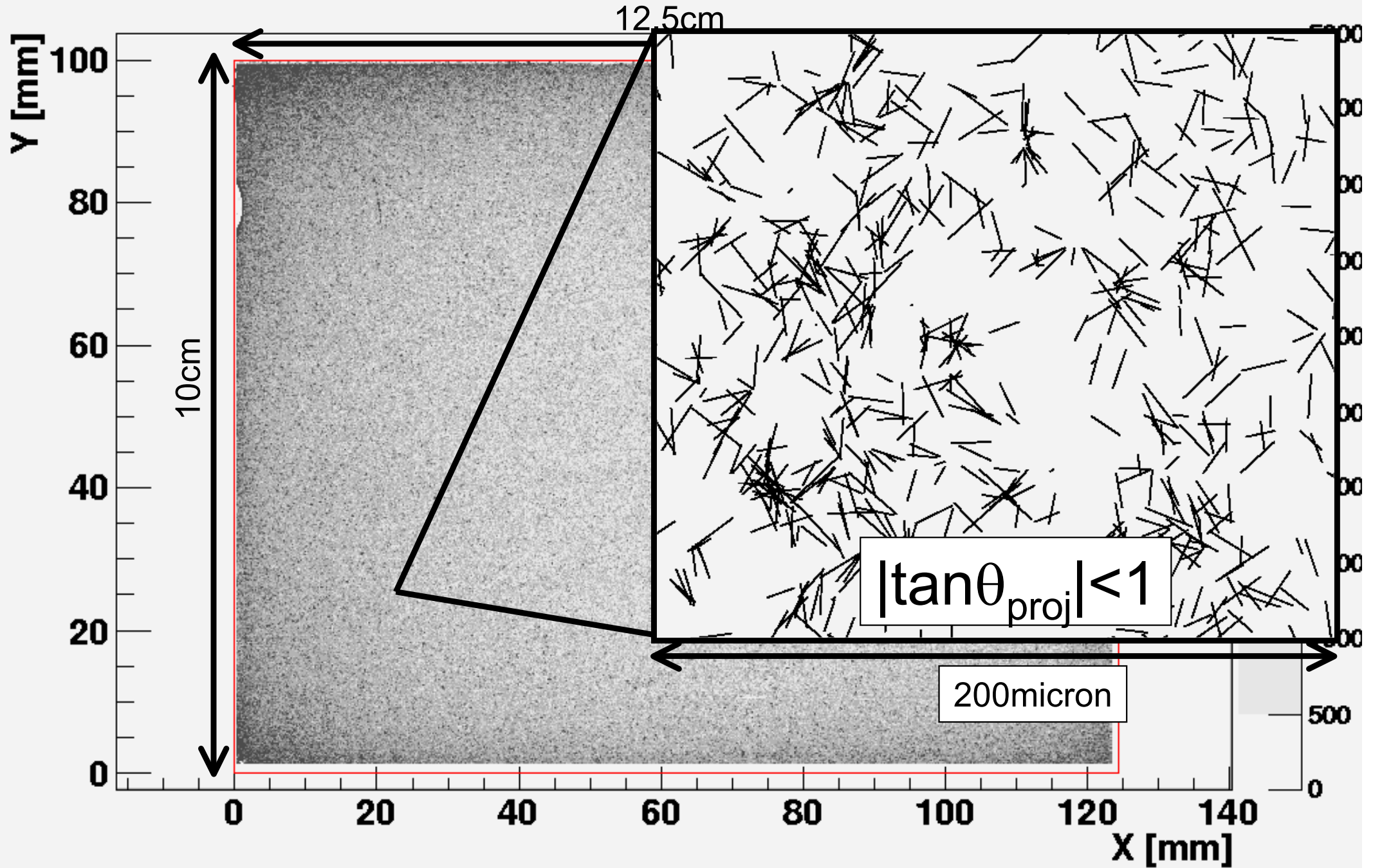
(OPERA film + lead plate(1mm))x17



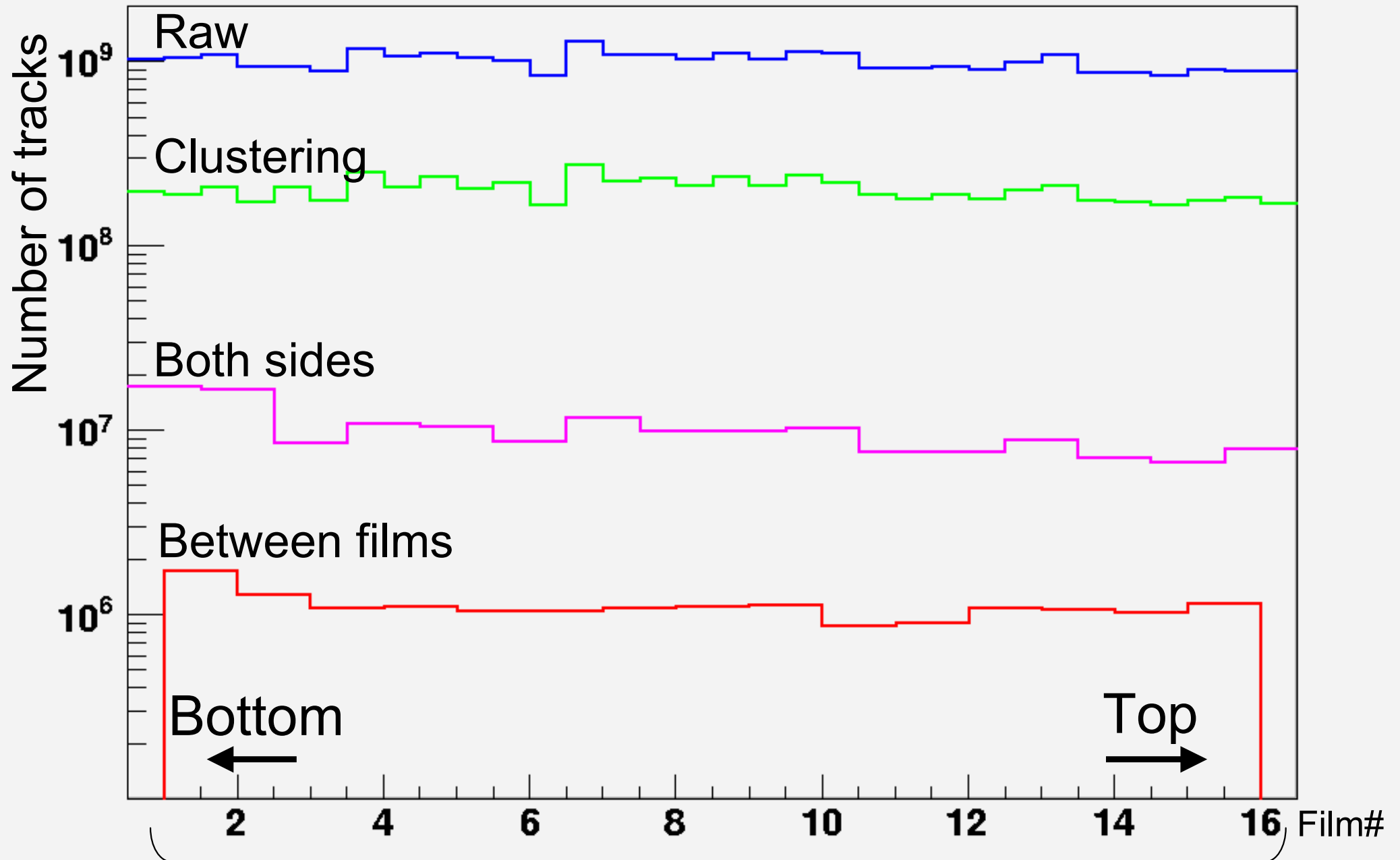
Number of tracks: 8.0×10^8



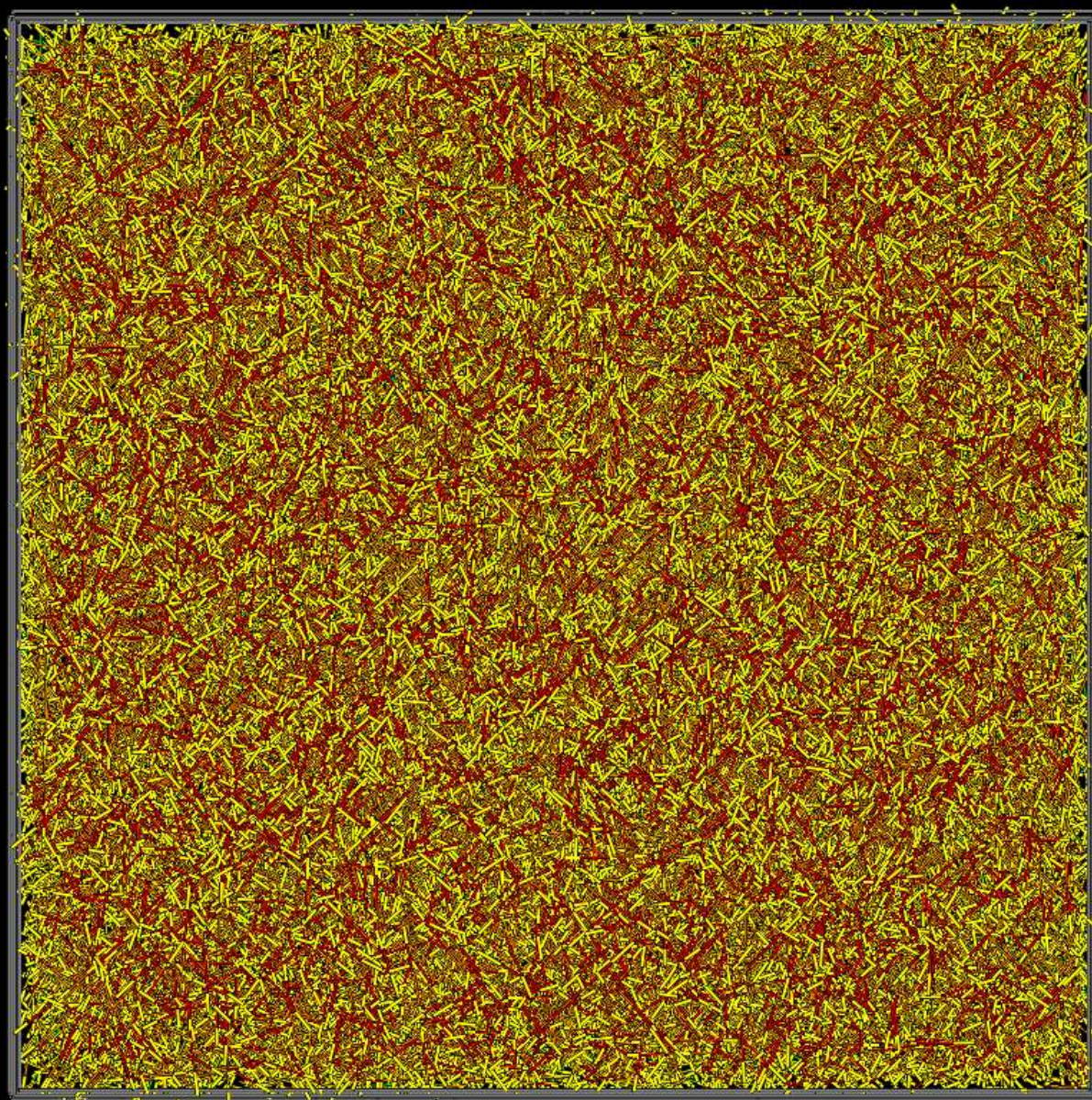
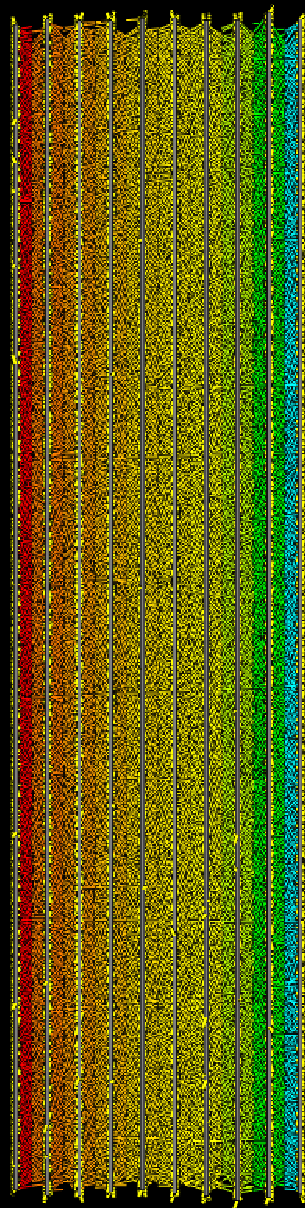
Number of tracks: 8.0×10^8



Track reconstruction

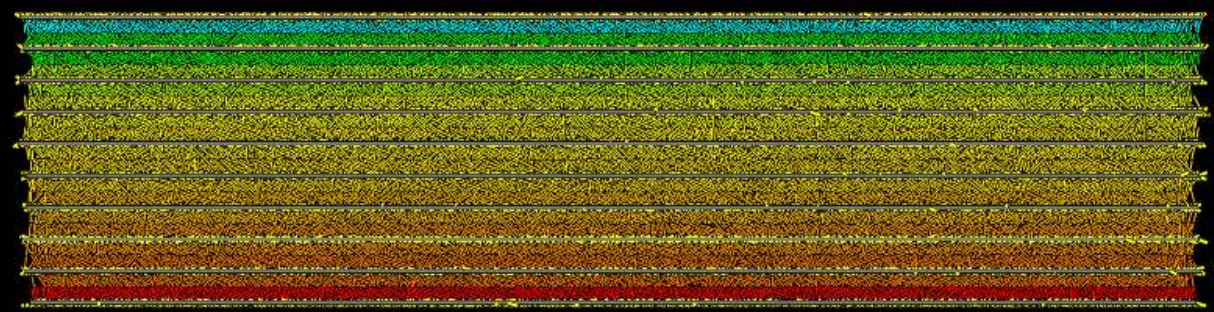


Reconstructed tracks (film#1-16) : 8.7×10^7 (overlap)



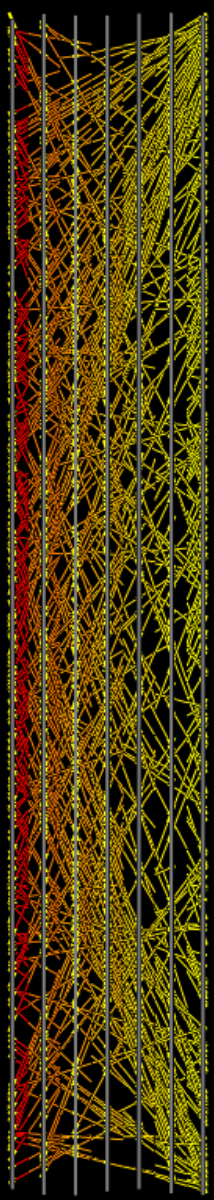
Z

Y

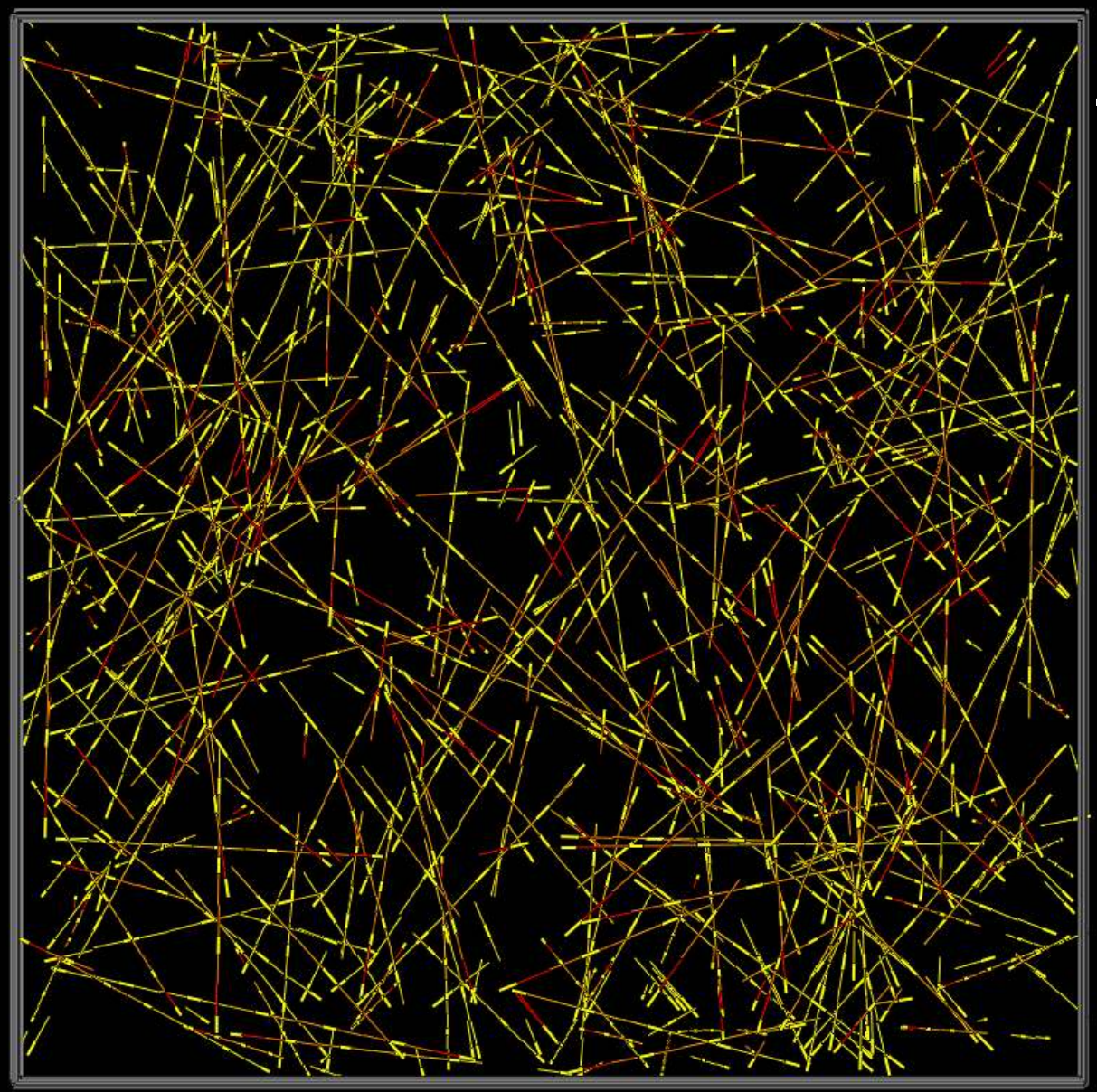


X

$(5\text{mm})^2 \times 10\text{films}$
 $2 \times 10^4\text{tracks}$

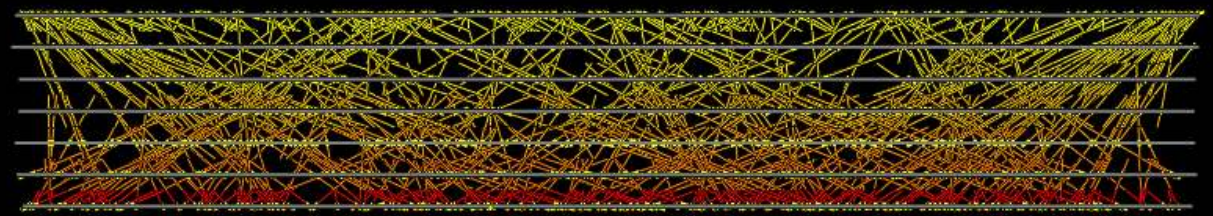


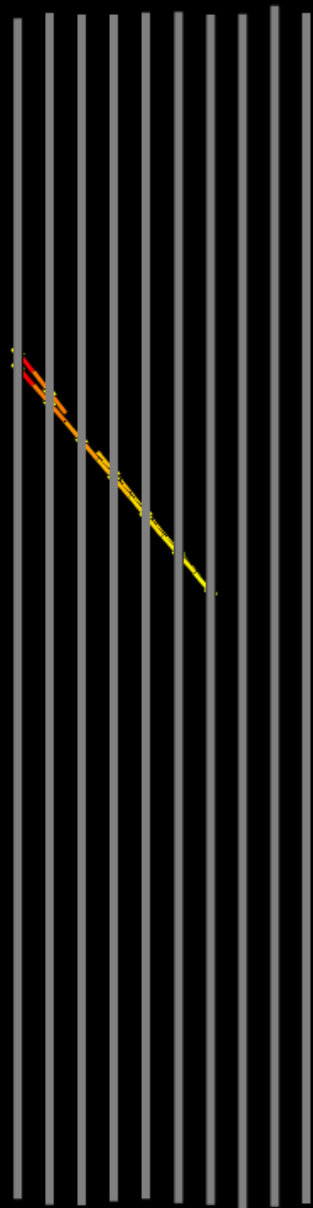
Y



Z

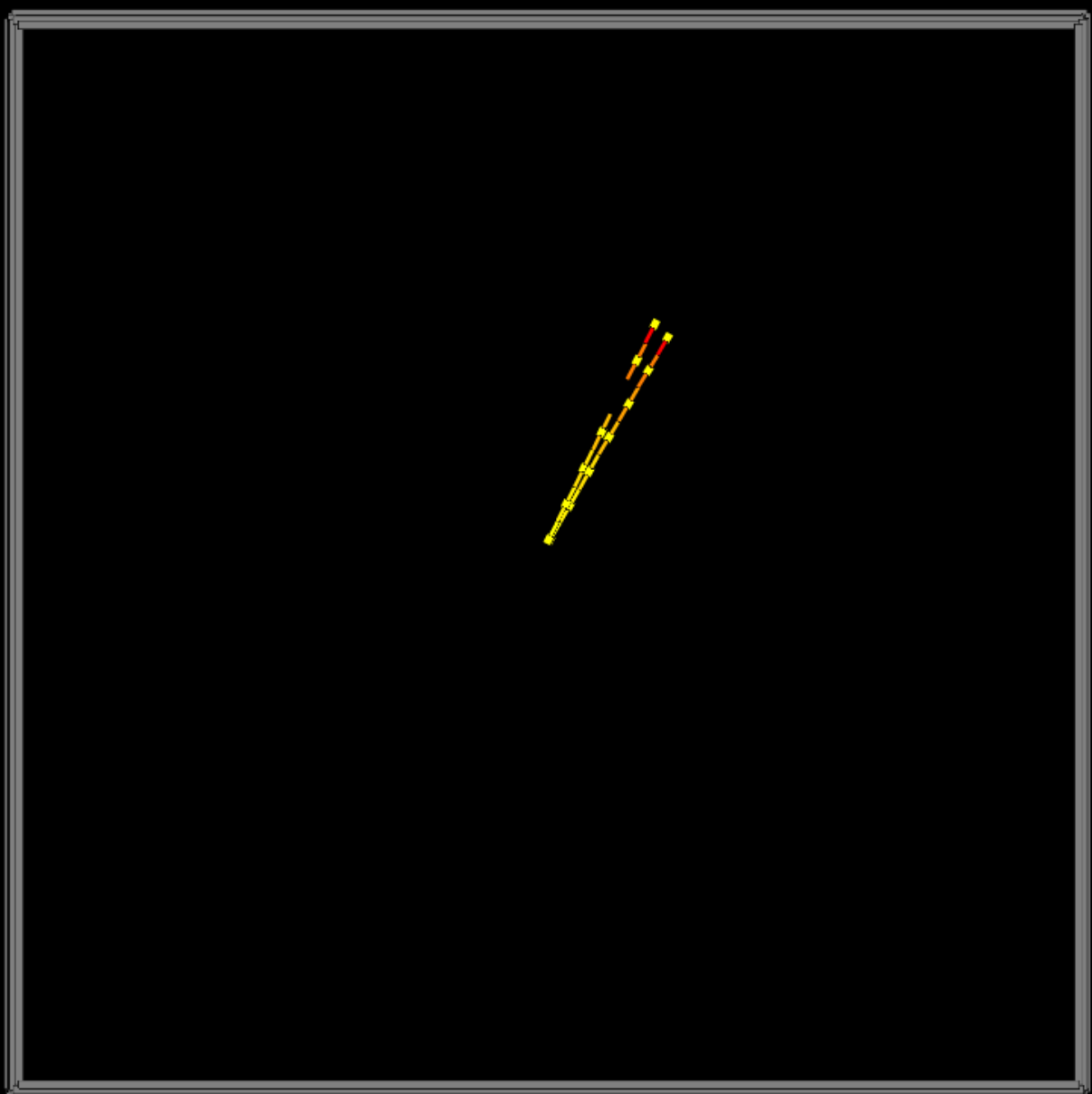
X



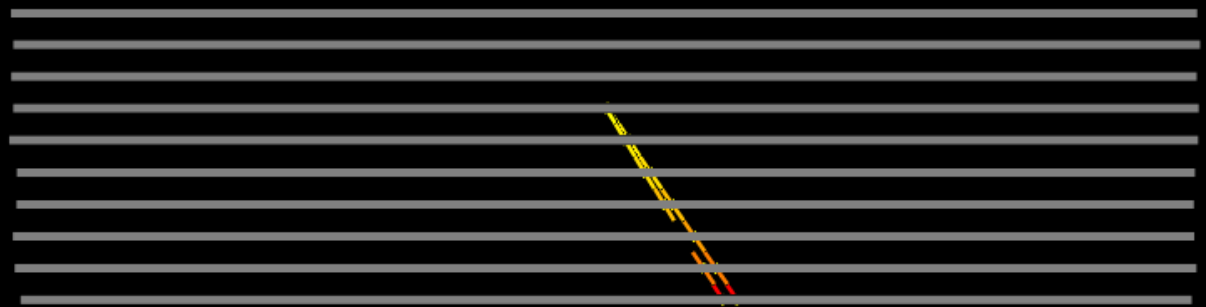


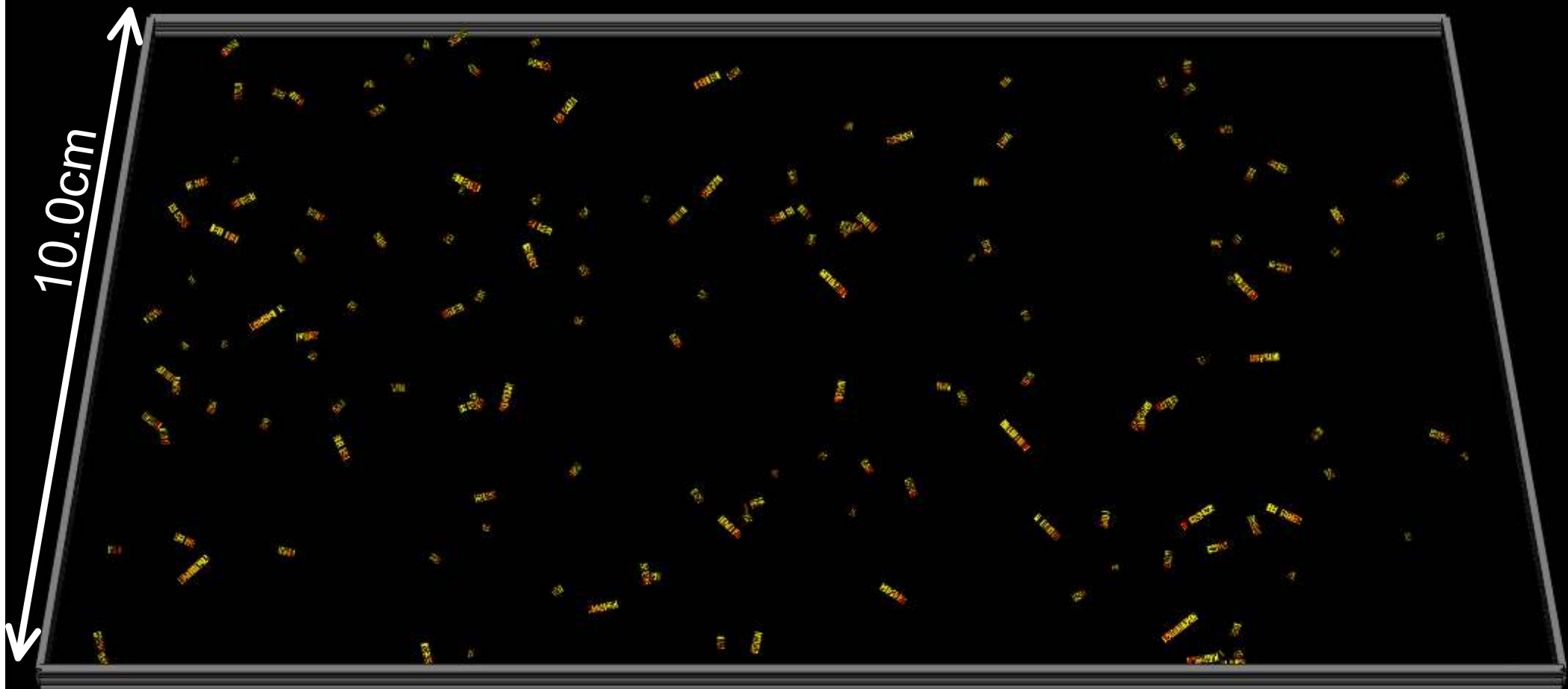
Y

X



Z



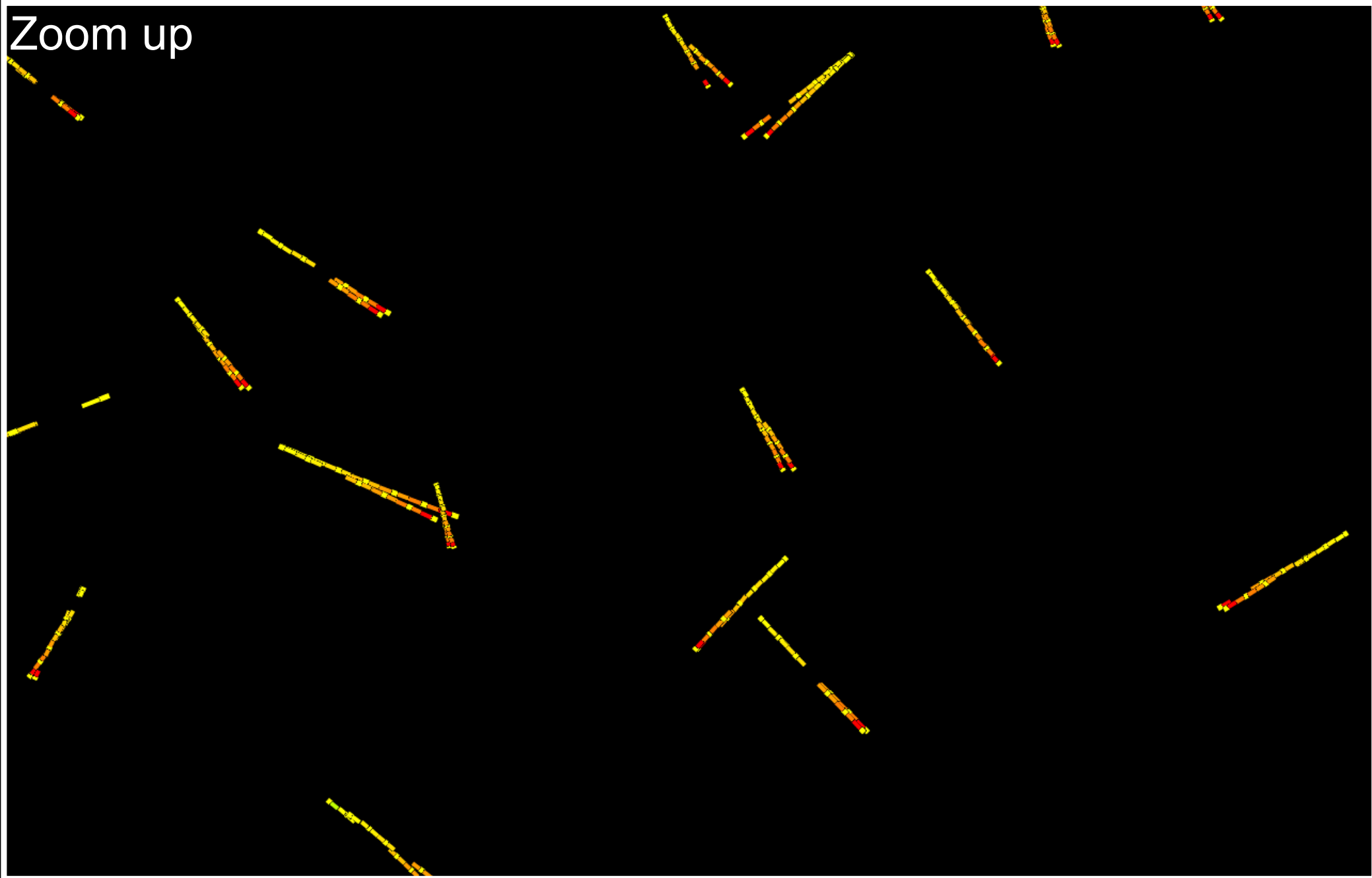


10.0cm

12.5cm

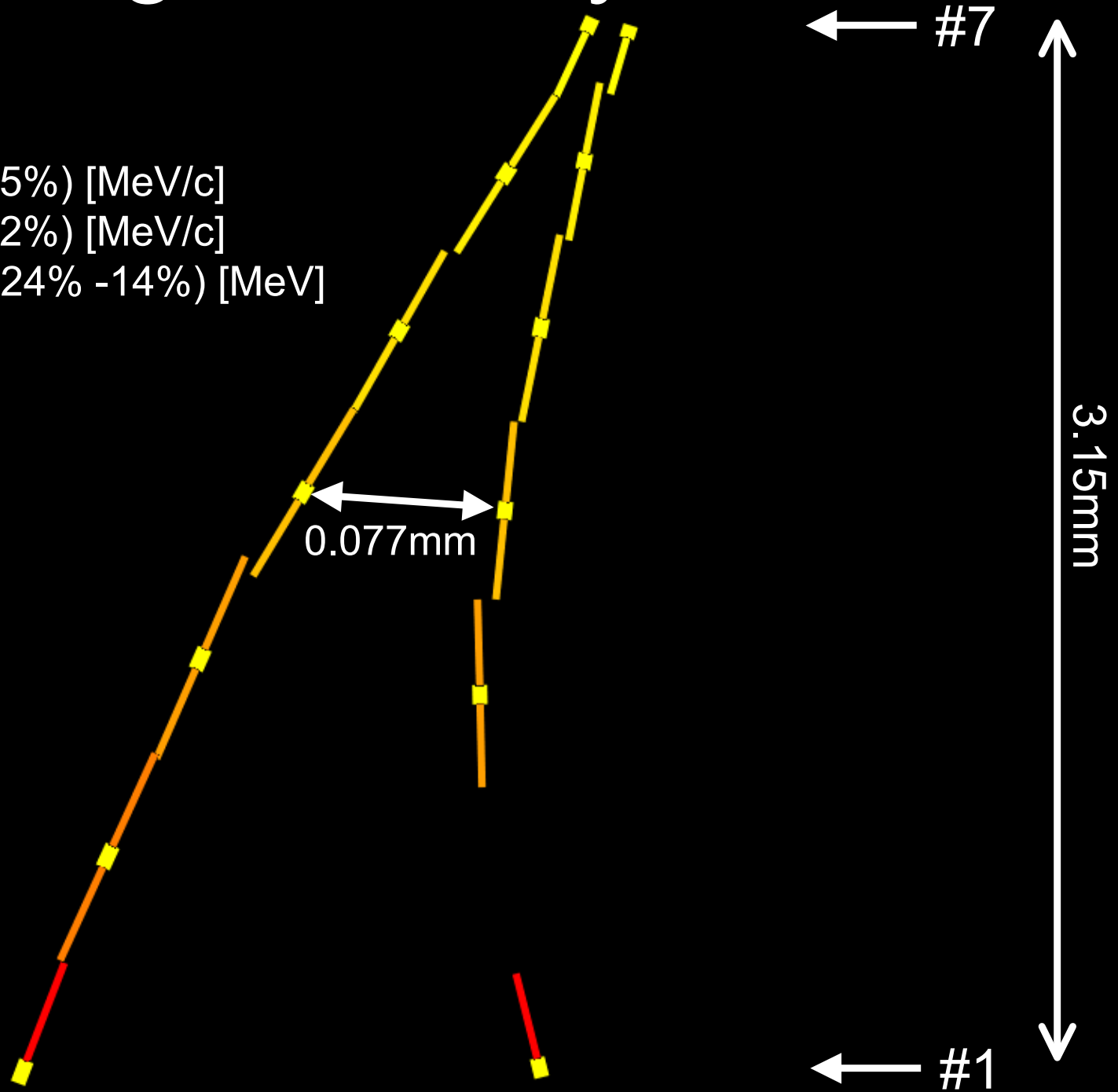
153events
Reliability 97%

Zoom up



One of gamma-ray events

Event : 71 6923485
Start : #7
 θ_{incident} : 9.748 [deg]
 $(p\beta)_{\text{left}}$: 60 +20 -12 (25%) [MeV/c]
 $(p\beta)_{\text{right}}$: 32 + 9 - 6 (22%) [MeV/c]
 E_{γ} : 92 +22 -13 (+24% -14%) [MeV]



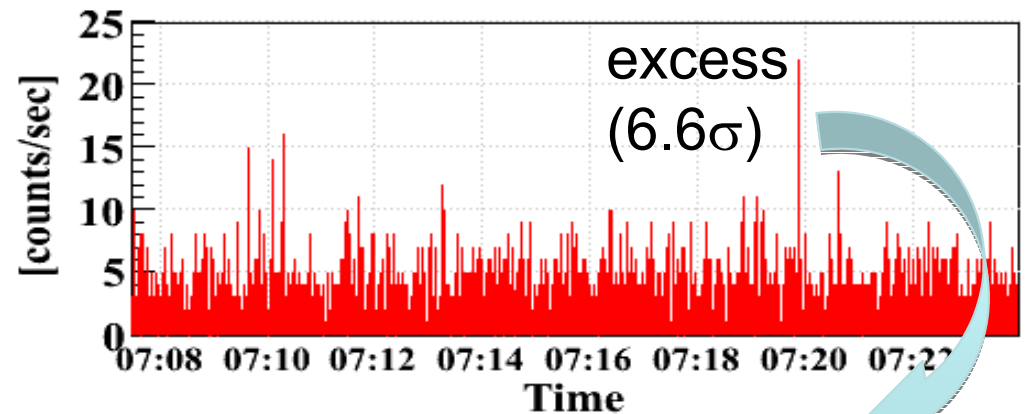
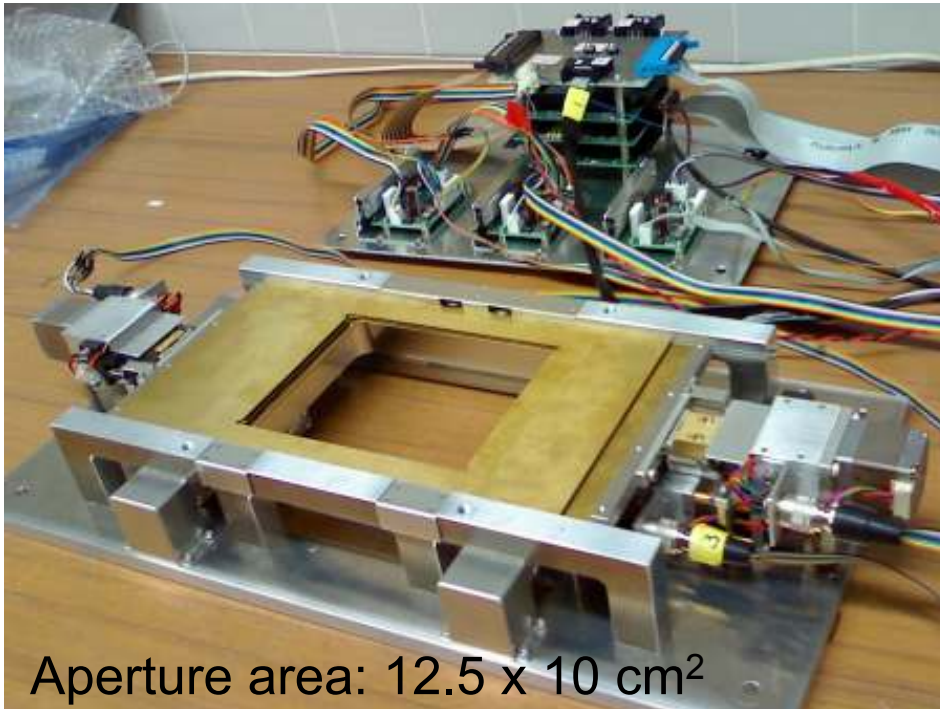
Establishment of timestamp technique

H.Rokujo, et al., NIM A, 701 (2013)

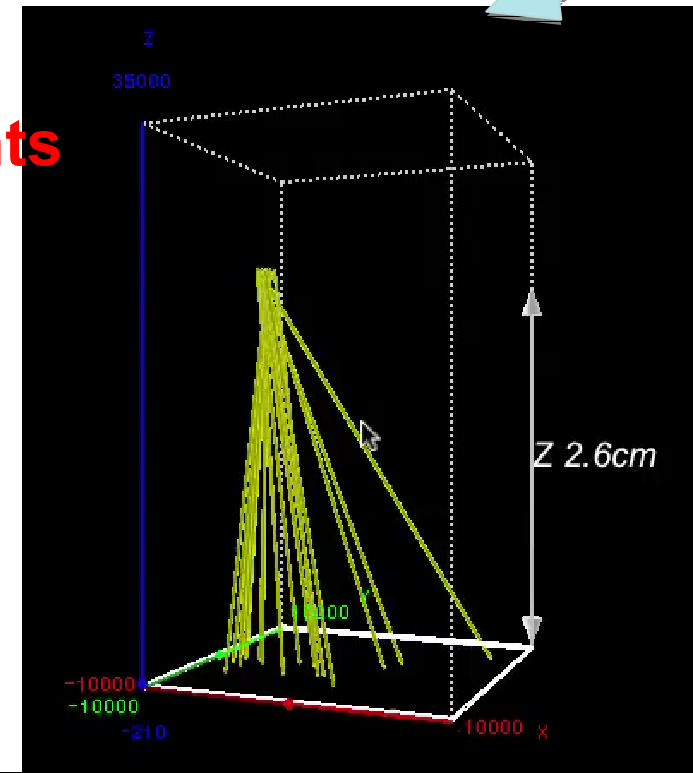
@GRAINE2011

“Multi-stage shifter” 1st model

Track rate measurement@35km



Detection of hadron events



- Correct operation during whole observation time
- Giving time info. to all penetrating tracks
- Detection of hadron shower tracks by timing and 3-D spatial analysis
- Time resolution: 0.15 sec

Hadron induced event

E_{ν} : 2438038

E_{γ} : 45^{+33}_{-10} [MeV]

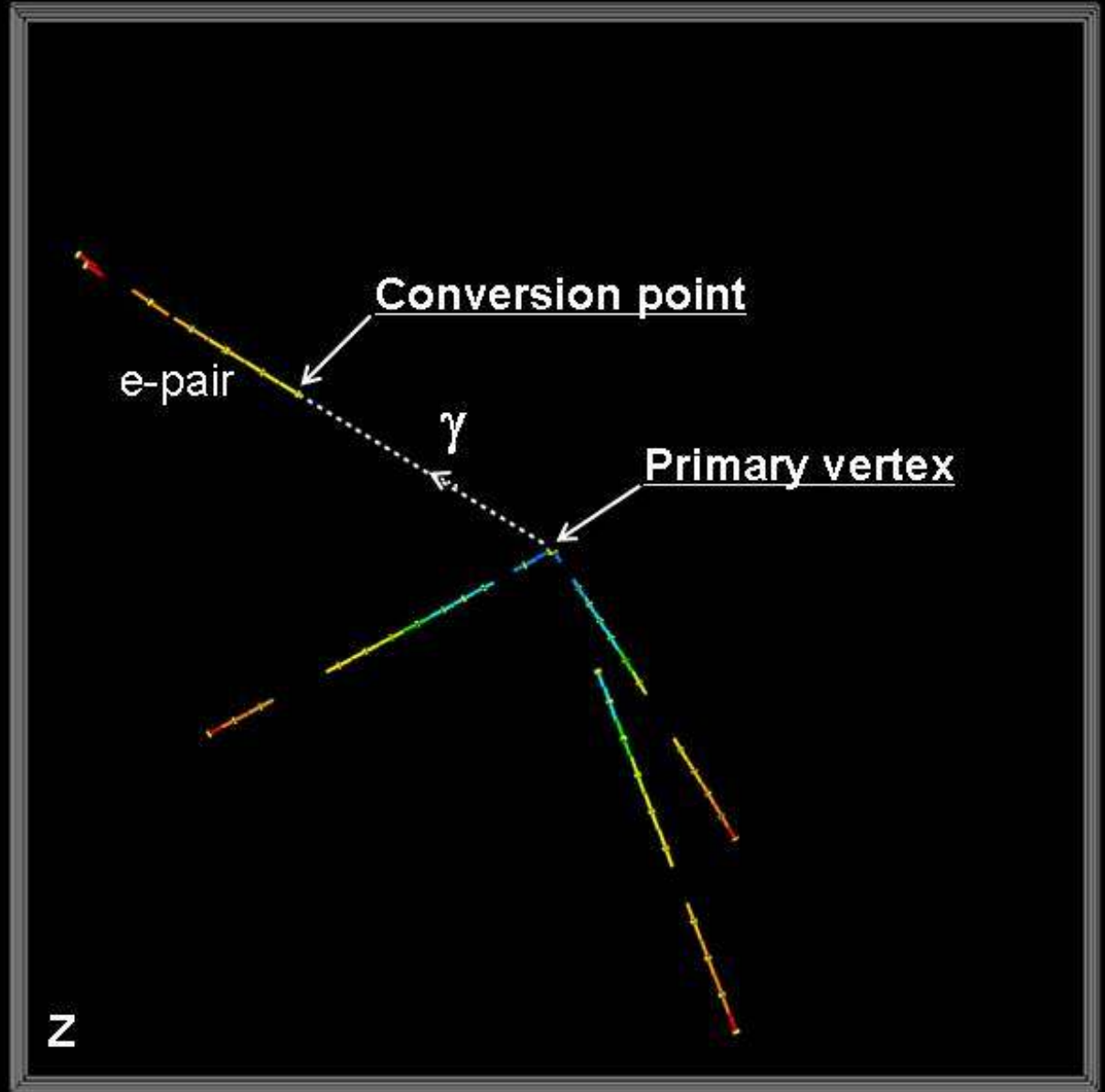
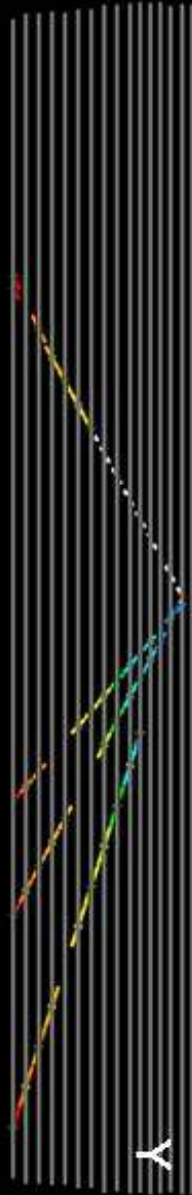
θ_{γ} : 46.61[deg]

7:18:34.5 (JST)

$\Delta t = \pm 0.5$ s

Convergence

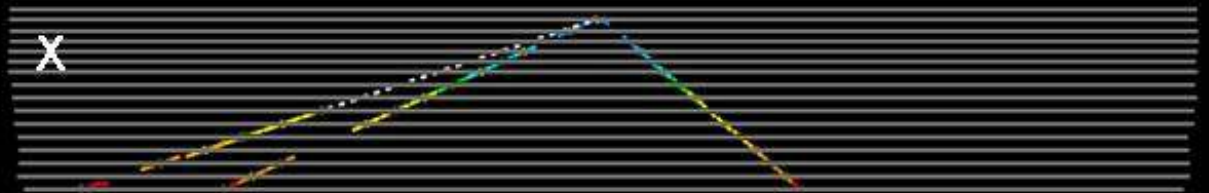
1.2cm x 1.2cm
x 16films



Pointing accuracy

$\Delta\theta_{\text{space}}$: 0.65deg

(0.0114rad)



Electron induced event

E_ν : 7797344

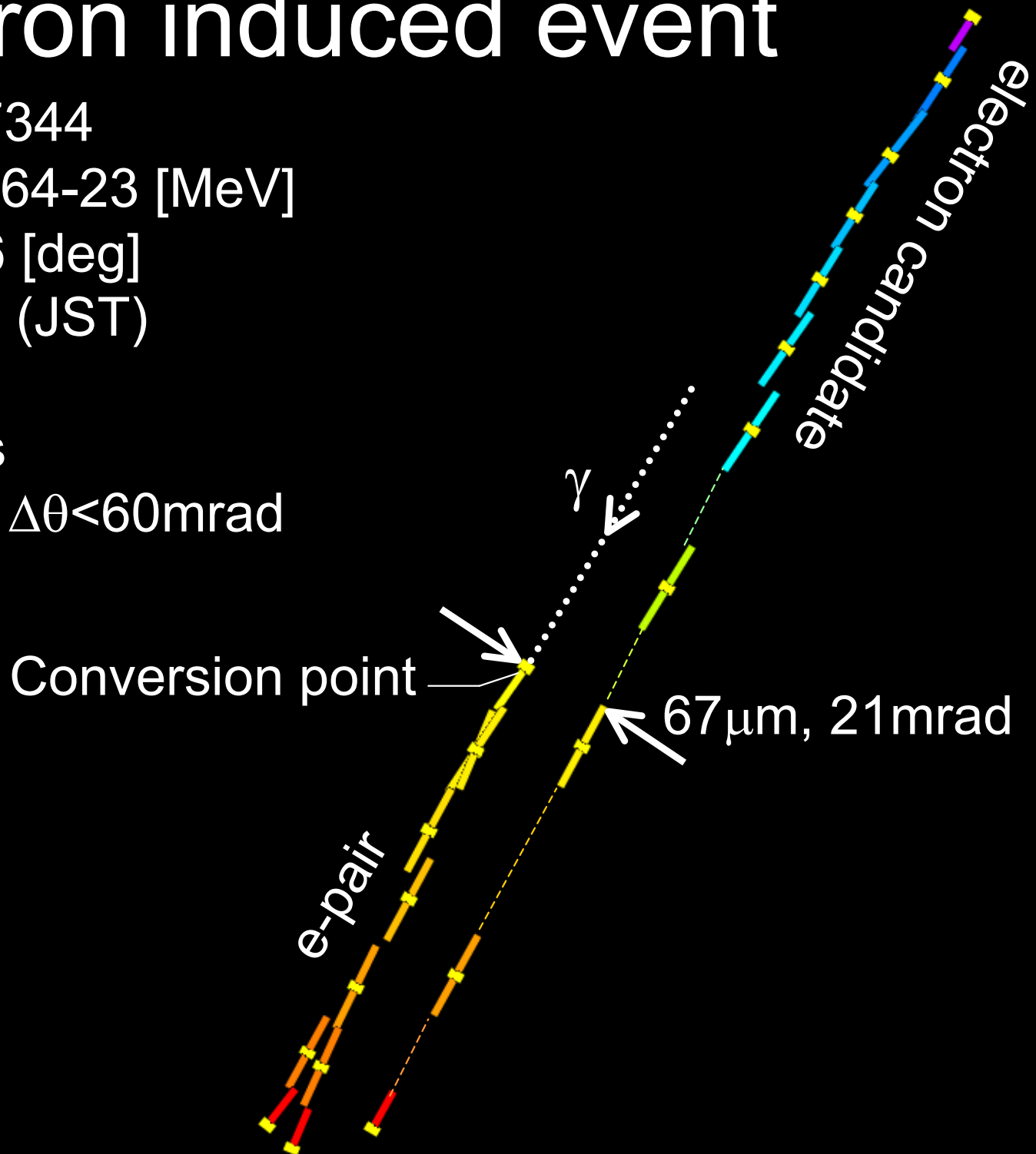
E_γ : 100+64-23 [MeV]

θ_γ : 38.96 [deg]

8:15:52.9 (JST)

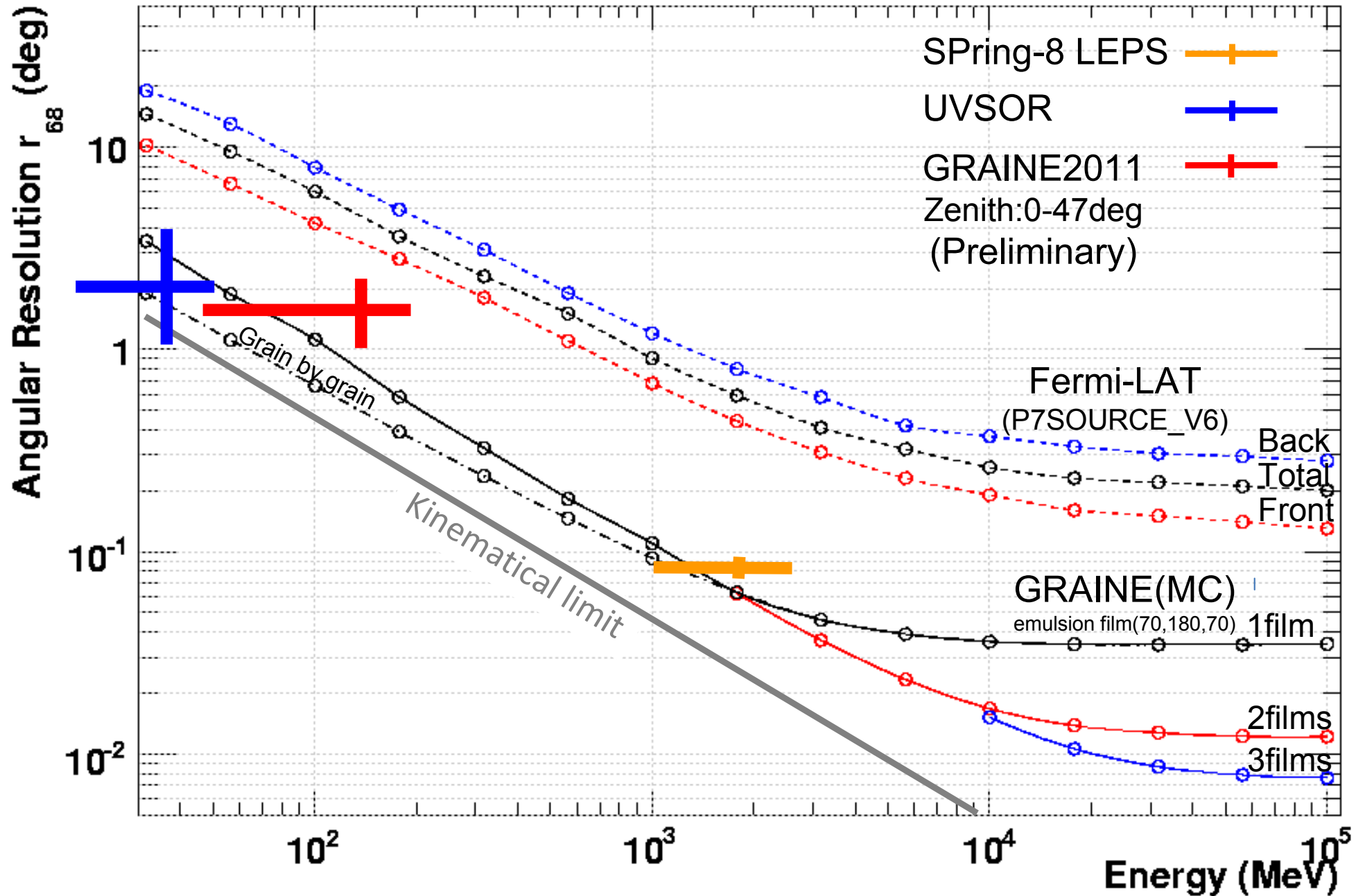
$\Delta t = \pm 0.5$ s

$\Delta r < 1$ mm, $\Delta\theta < 60$ mrad

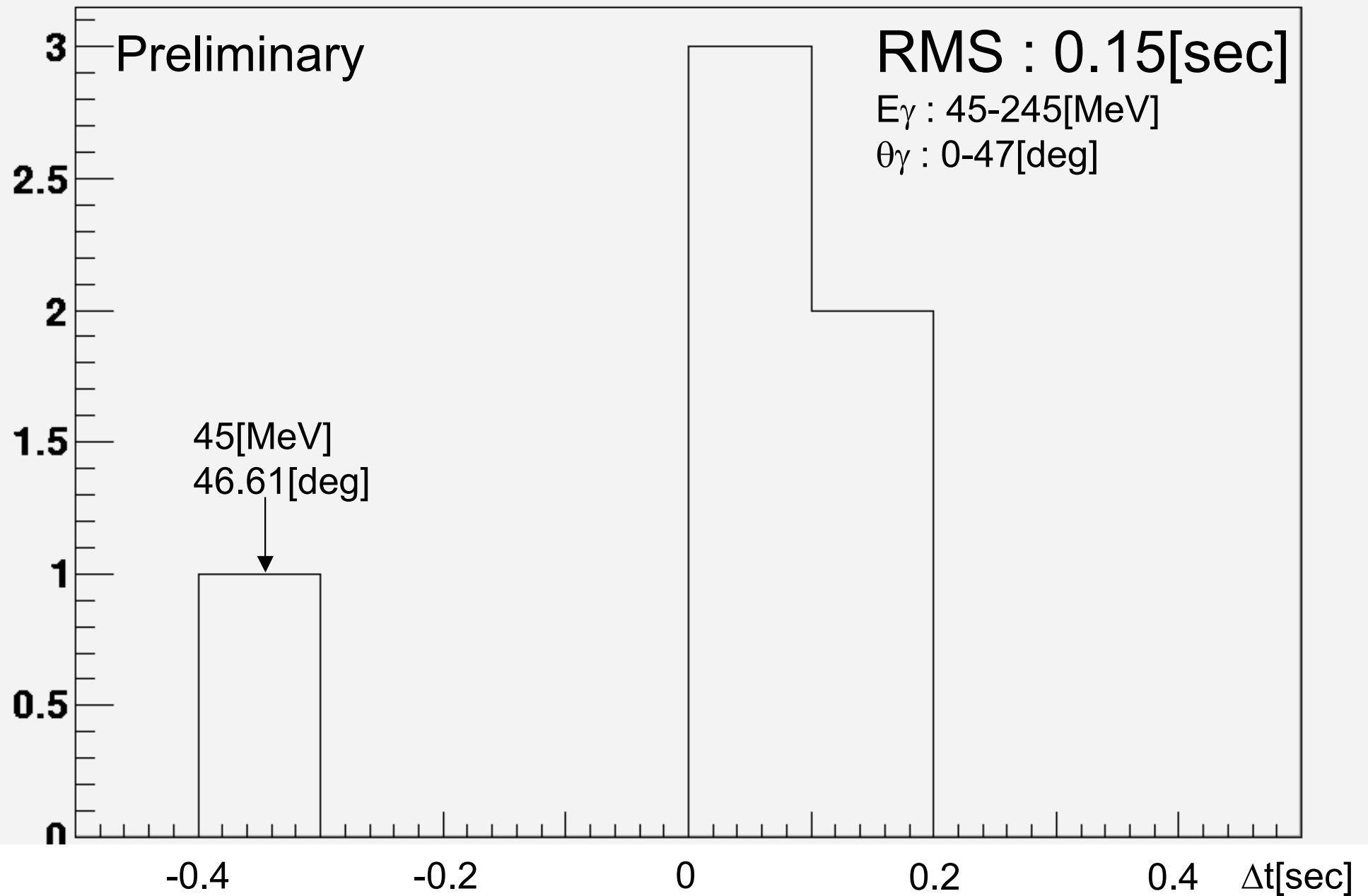


Angular resolution

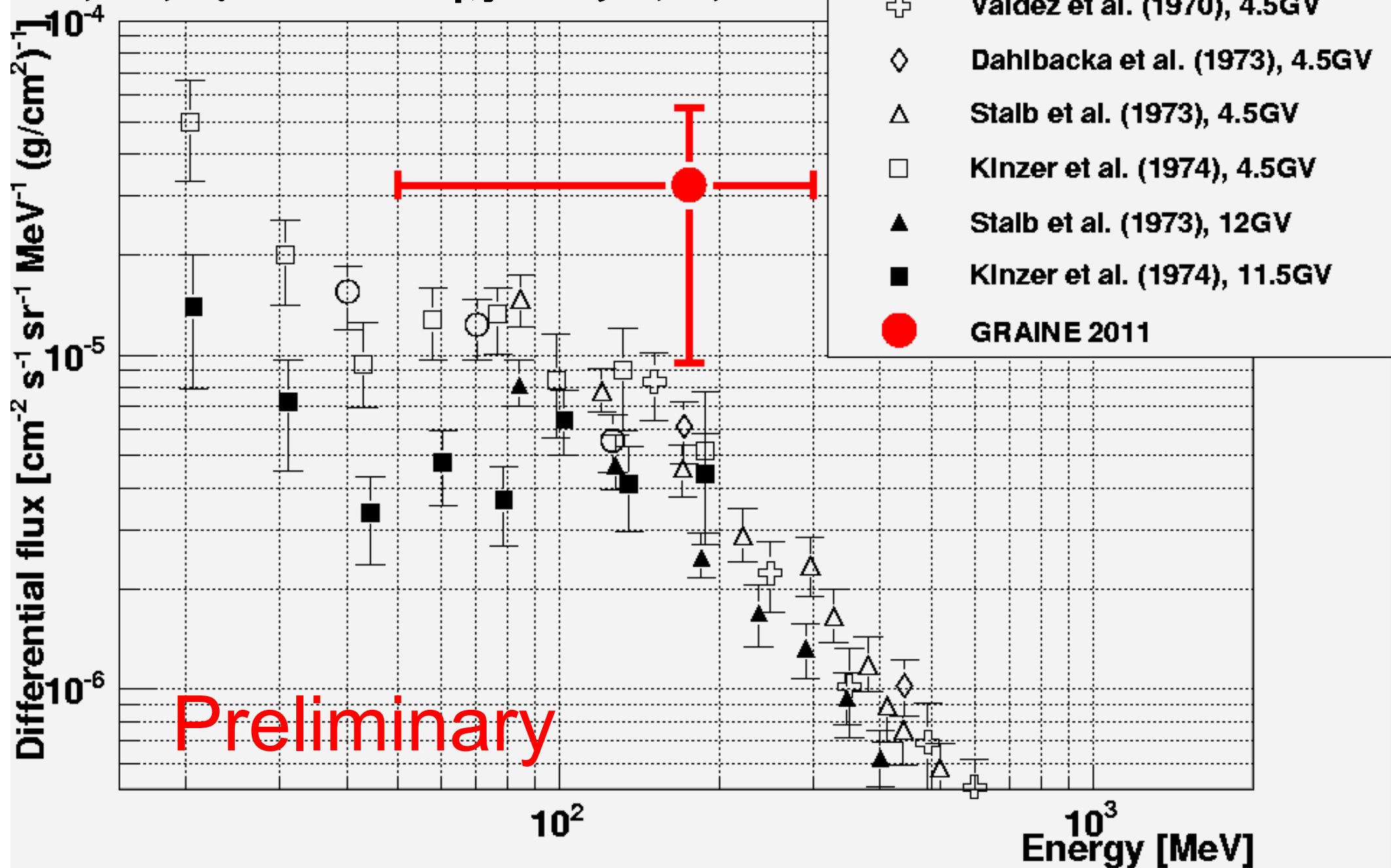
PSF at normal incidence



Gamma-ray timing accuracy



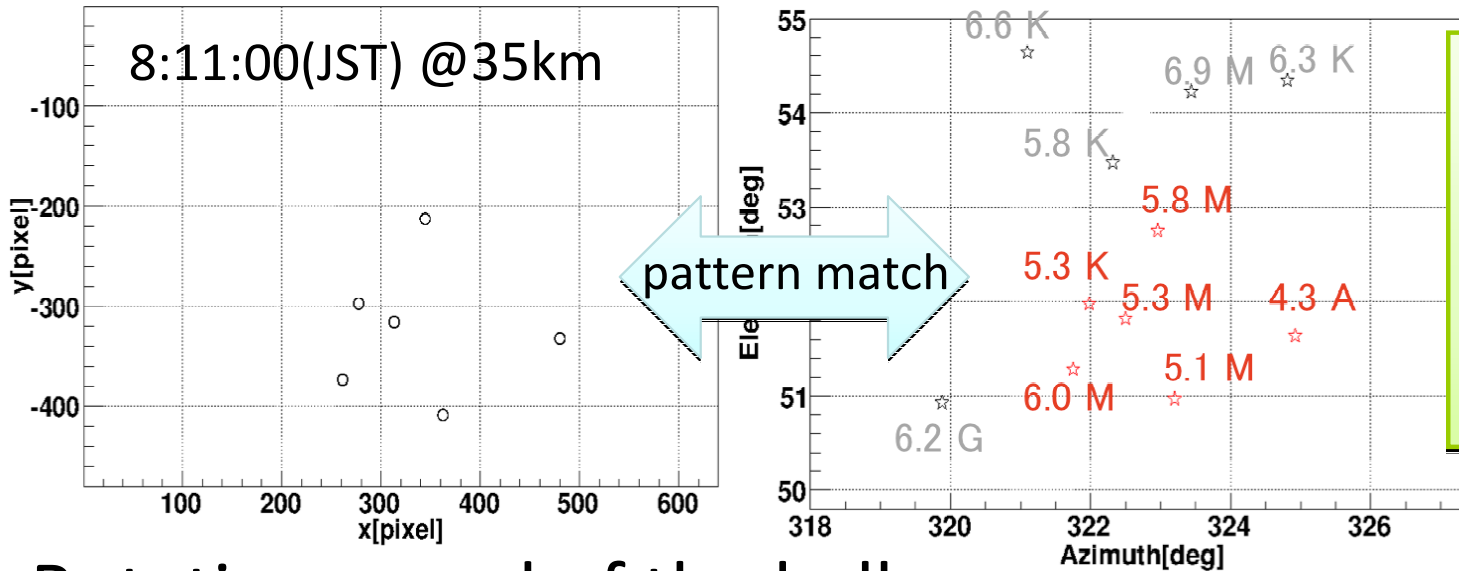
大気ガンマ線フラックス



Attitude analysis

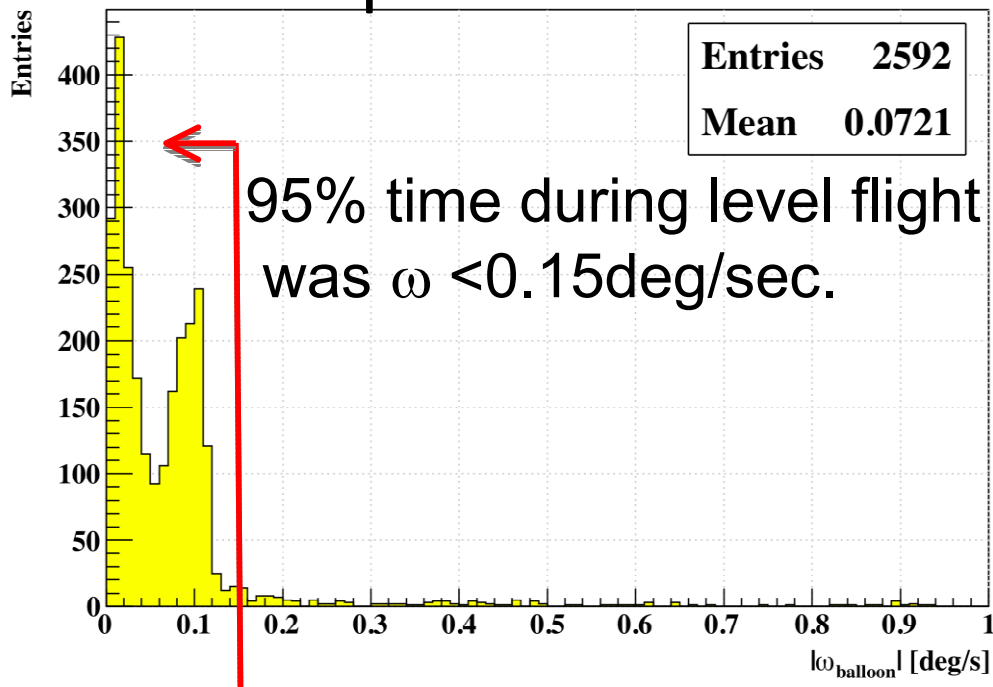
Daytime star camera view

Star catalog data



Working rate: 74 %
Monitoring
accuracy: < mrad
Elevation < 0.25mrad
Azimuth < 0.44mrad

Rotation speed of the balloon



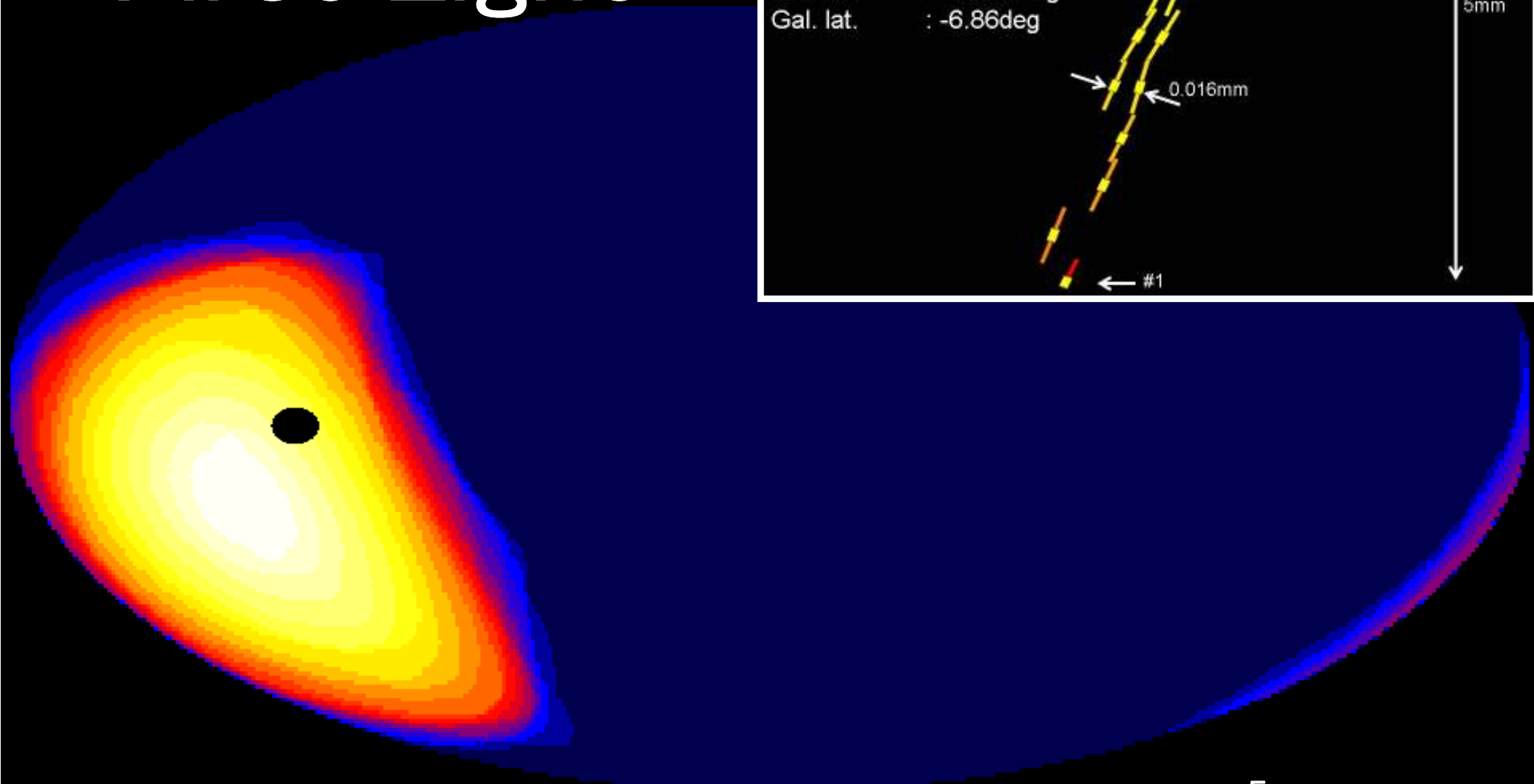
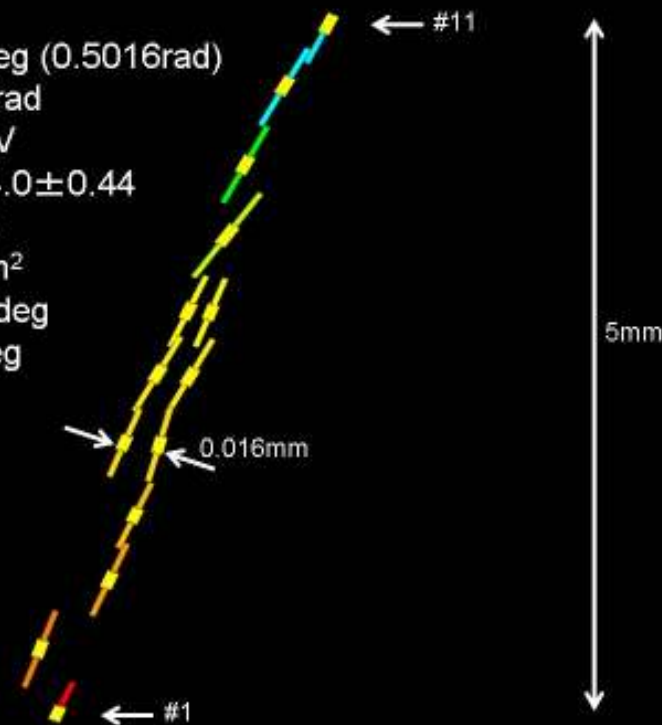
It is important to decide telescope attitude to celestial coordinate better than emulsion angular resolution(0.08deg).

**We confirmed attitude decision accuracy was $< \omega \sigma_t$
< 0.02deg.**

GRAINE

First Light

Event : 111 2986322
Start : #11 up
 θ_{incident} : 26.64deg (0.5016rad)
 θ_{open} : 0.0059rad
 $E_{\gamma}(\theta_{\text{open}})$: 340MeV
JST : 8:24:44.0 \pm 0.44
Altitude : 34.6km
Atm. depth : 6.6g/cm²
Gal. lon. : 112.06deg
Gal. lat. : -6.86deg

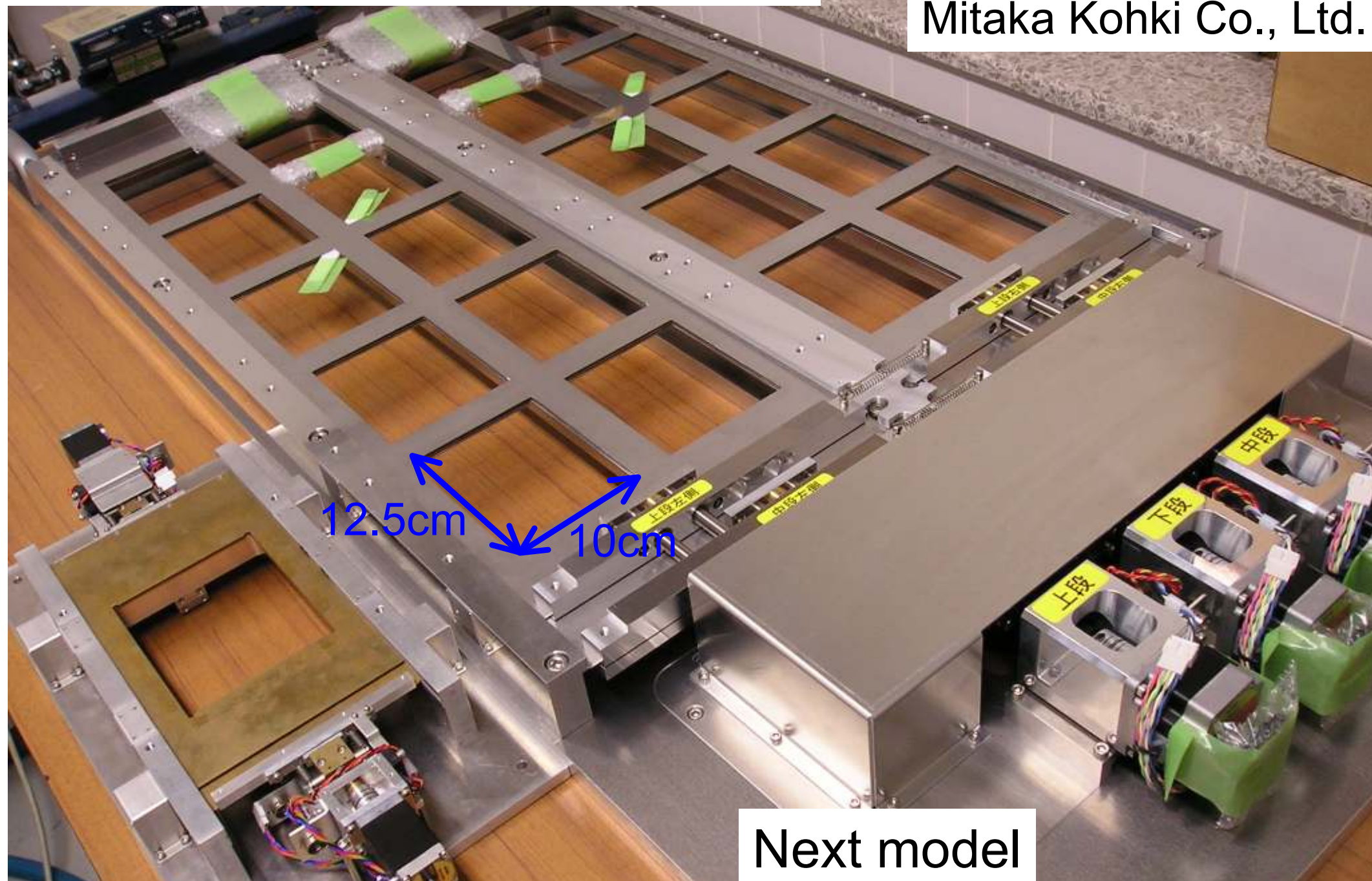


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Flight model of multi-stage shifter

Co-developed with
Mitaka Kohki Co., Ltd.



12.5cm
10cm

Next model

2011 model

Aperture area : 2500cm²