

Cherenkov Telescope Array Observatory (CTAO) 計画：全体報告（27）



齋藤隆之、窪秀利、手嶋政廣、戸谷友則、野崎誠也、吉越貴紀(東京大)、野田浩司(千葉大)、吉田龍生(茨城大)、井岡邦仁(京都大)、田島宏康(名古屋大)、山本常夏(甲南大)
他CTA Consortium

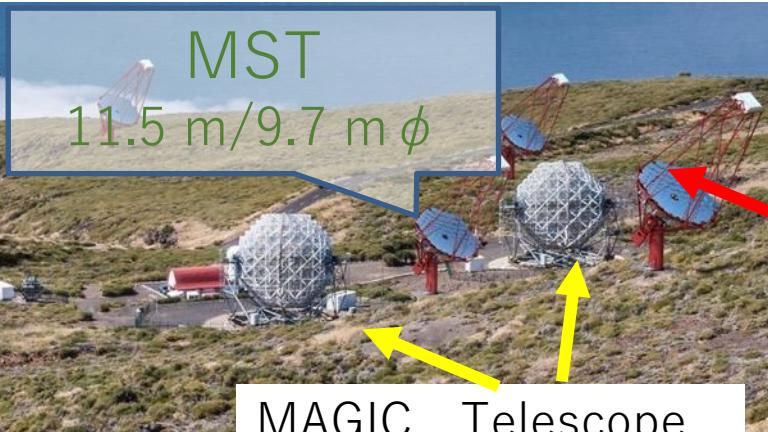
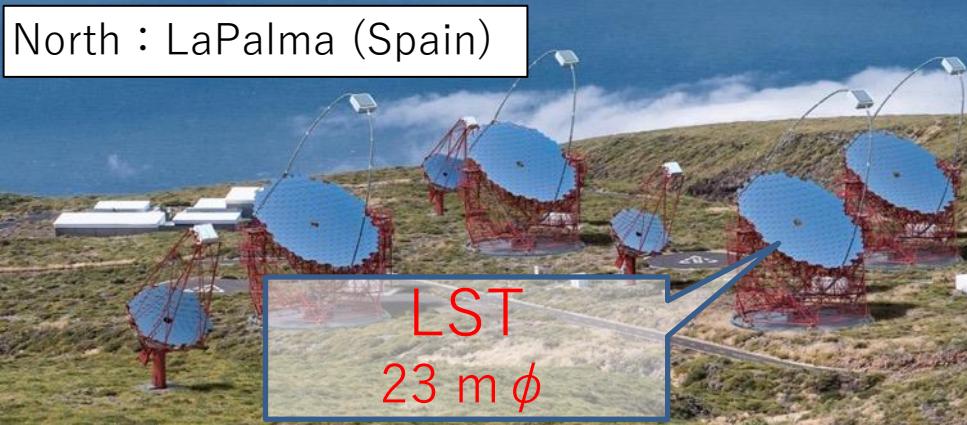
25か国
 >1500名



CTA-Japan 127名

青山大	大林花織, 佐藤優理, 田中周太, 山崎了, 吉田篤正	東大宇宙線研	浅野勝晃, 阿部正太郎, 粟井恭輔, 糸川拓海, 猪目祐介, 大石理子, 大岡秀行, 大谷恵生, 窪秀利, 斎藤隆之, 武石隆治, 手嶋政廣, 戸村友宣, 野崎誠也, バクスター ジョシュア 稜, 橋山和明, 吉越貴紀, Y. Chai, D. Hadasch, D. Mazin, M. Strzys, I. Vovk, P. K. H. Yeung
茨城大	片桐秀明, 柳田昭平, 吉田龍生		共同研究員 : 岡知彦, 櫻井駿介, 広谷幸一, 深見哲志, 村瀬孔大, K. S. Cheng, X. Cui, D. C. Y. Hui, A. K. Kong, P. Majumdar, J. Takata, T. P. H. Tam, W. Tian, L. Wan
宇宙研	鈴木 寛大, 林克洋		
大阪大	井上芳幸, 藤原立樹, 松本浩典		
神奈川大	辻直美		
北里大	村石浩	東北大	當真賢二, 石崎涉
岐阜大	佐野栄俊	徳島大	折戸玲子
九州大	稻田知大	名大理	立原研悟, 早川貴敬, 福井康雄, 山本宏昭
京大基研	井岡邦仁	名大ISEE	安藤大地, 奥村暁, 重谷優斗, 河原崎琉, 高橋光成, 田島宏康, S. Bang
京大理	鶴剛, 寺内健太, 李兆衡		
熊本大	高橋慶太郎	広大先理工	今澤遼, 樞木大修, 木坂将大, 須田祐介, 高橋弘充, 仲野悟帆, 橋爪大樹, 深沢泰司, A. Roy
KEK	田中真伸		
甲南大	井上剛志, 田中孝明, 千川道幸, 溝手雅也, 山本常夏	広大宇科セ	水野恒史
国立天文台	郡和範	宮崎大	森浩二
埼玉大	勝田哲, 清本拓人, 寺田幸功	山形大	郡司修一, 門叶冬樹, 中森健之
仙台高専	加賀谷美佳, 林航平	山梨学院大	内藤統也, 原敏
千葉大	井上進, 小林志鳳, 野田浩司	横浜国大	廣島渚
東海大	阿部和希, 韓天舒, 櫛田淳子, 佐藤雄輝, 西嶋恭司, 姚屹	理研	長瀧重博, M. Barkov, G. Ferrand, H. He, E. R. Owen, D. Warren
東大理	大平豊, 立石大, 戸谷友則, 馬場彩	立教大	内山泰伸, 澤田真理, 林田将明
東京都立大	川中宣太, 藤田裕	早稲田大	片岡淳

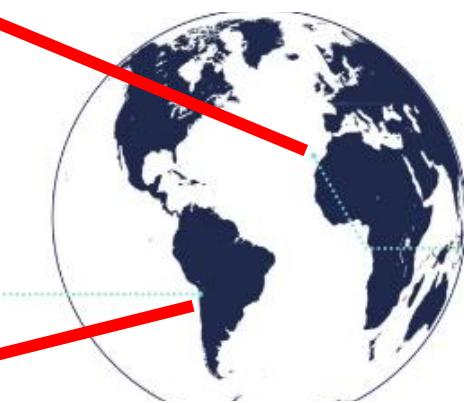
North : LaPalma (Spain)



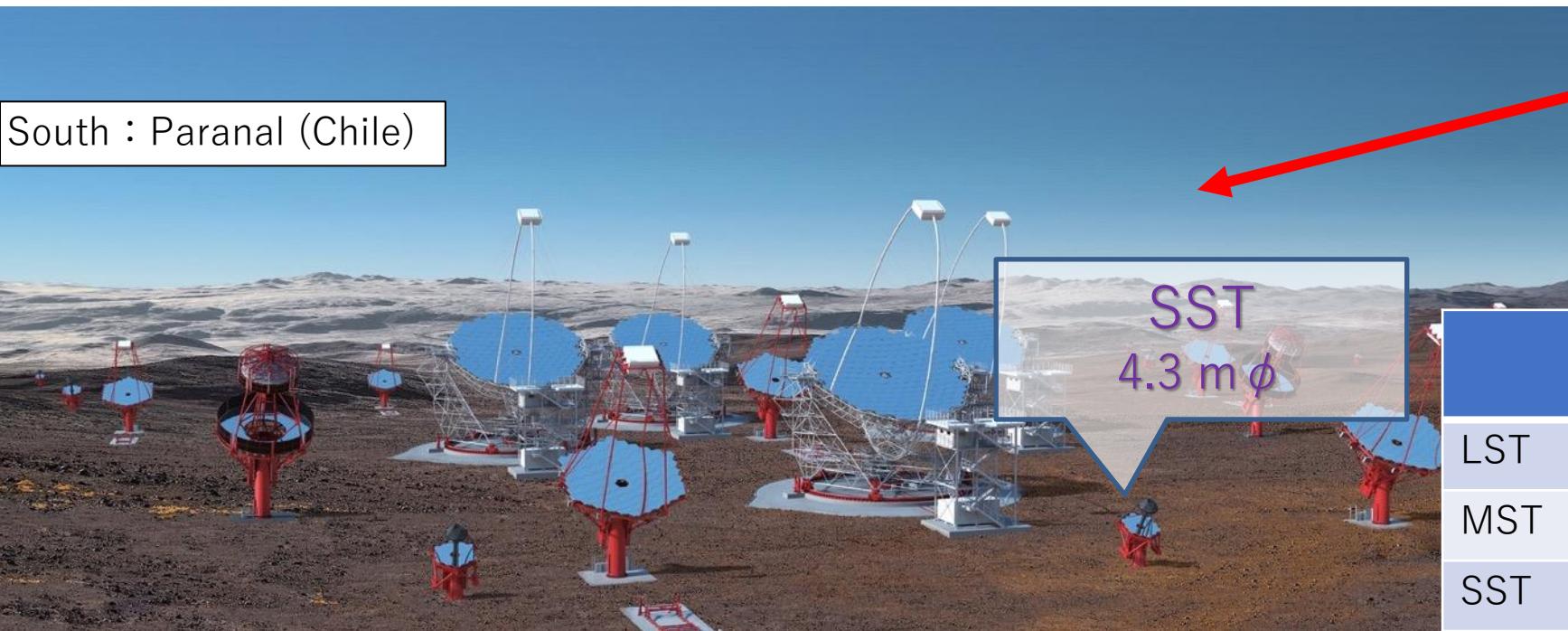
North

Array Coordinates

0.5 km x 0.5 km

Latitude: $28^{\circ} 45' 43.7904''$ North
Longitude: $17^{\circ} 53' 31.218''$ West

South : Paranal (Chile)



South

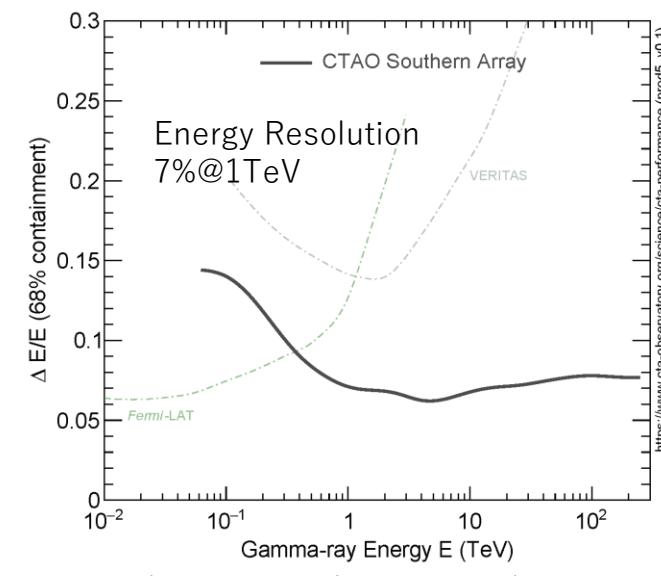
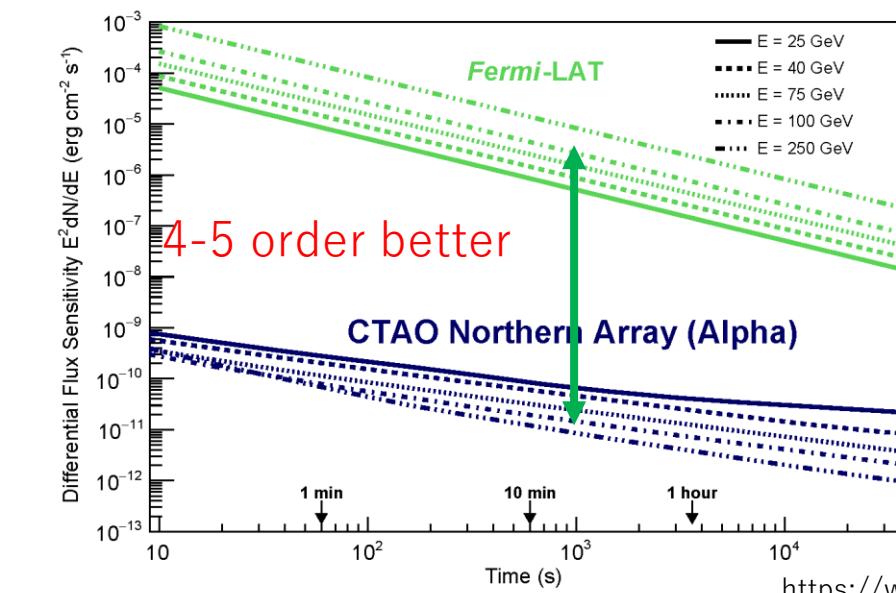
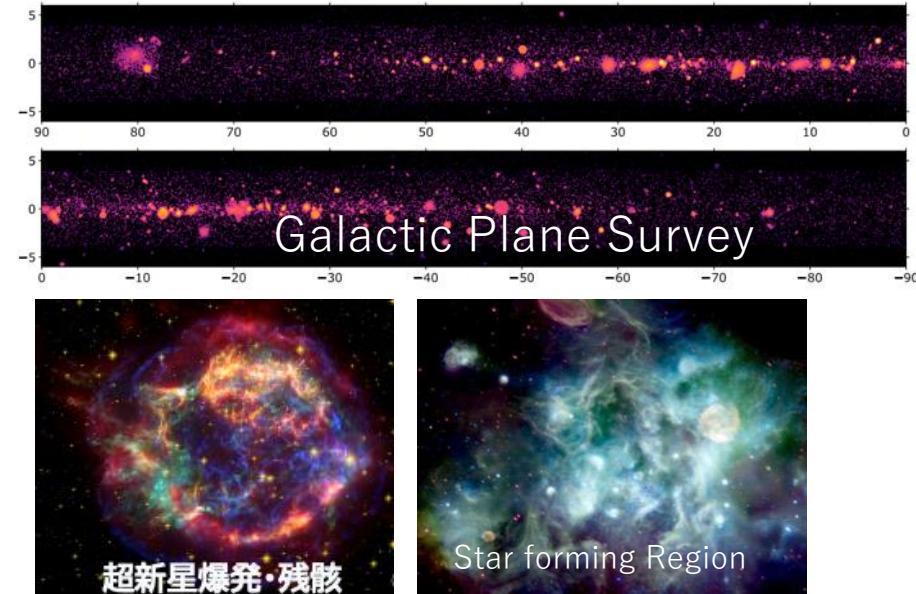
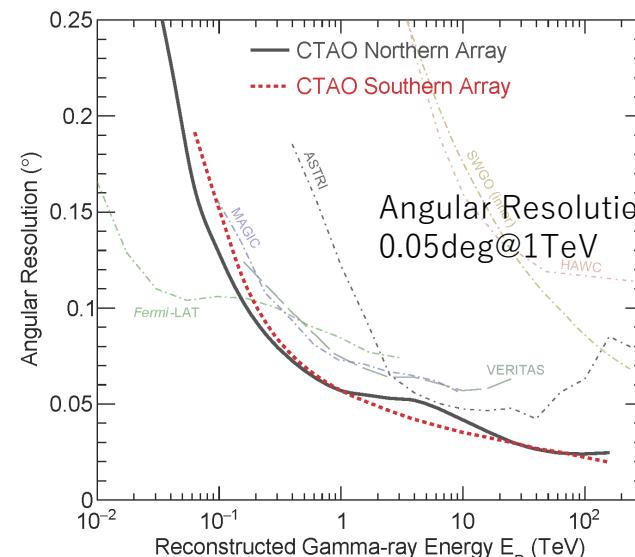
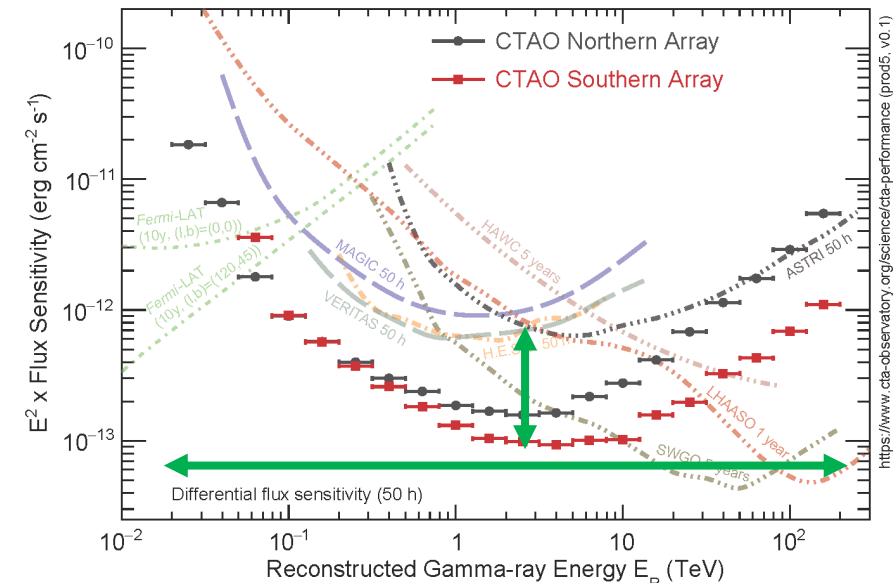
Array Coordinates

2 km x 2 km

Latitude: $24^{\circ} 41' 0.34''$ South
Longitude: $70^{\circ} 18' 58.84''$ West

	#tel (N)	#tel (S)	Energy [TeV]	FoV [deg]
LST	4	2→4	0.02 - 3	4.5
MST	9→15	14→25	0.08 - 50	7.5
SST	0	42→70	1 - 300	10.5

“Alpha” Configuration Operation starts in 2028.



LST-1



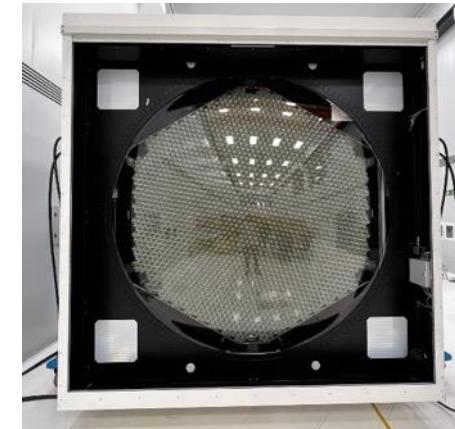
LST-3



LST-2



LST-4



Camera@Tenerife

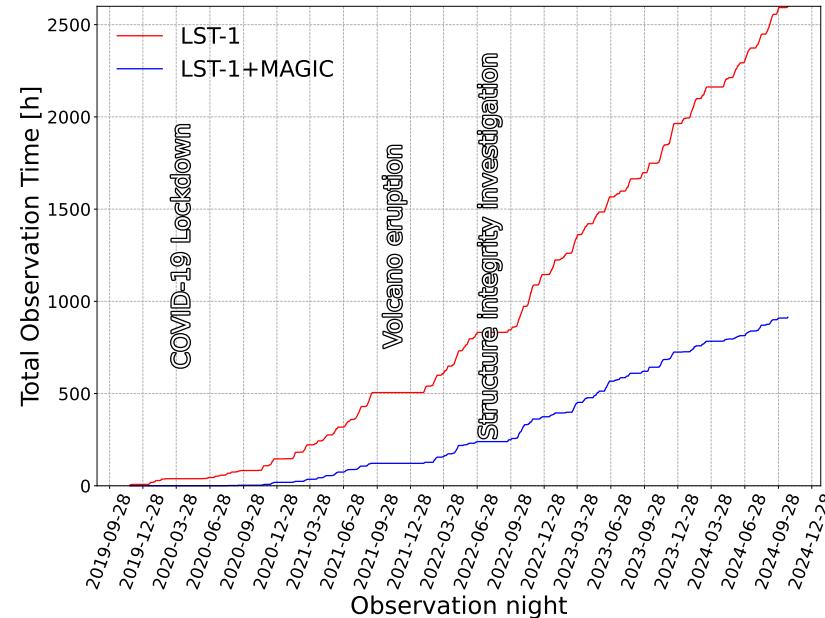
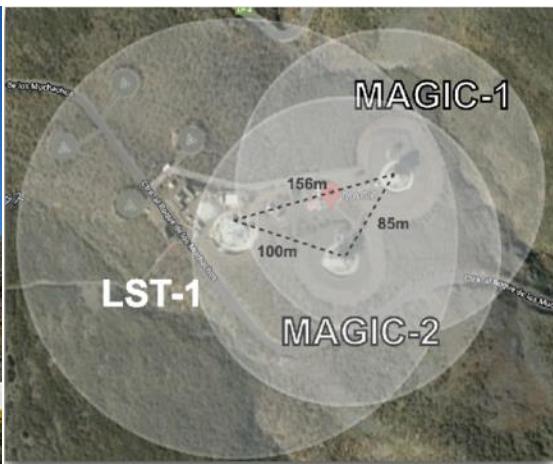
Schedule



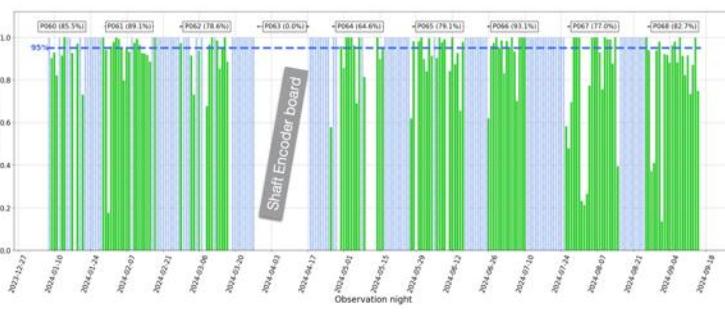
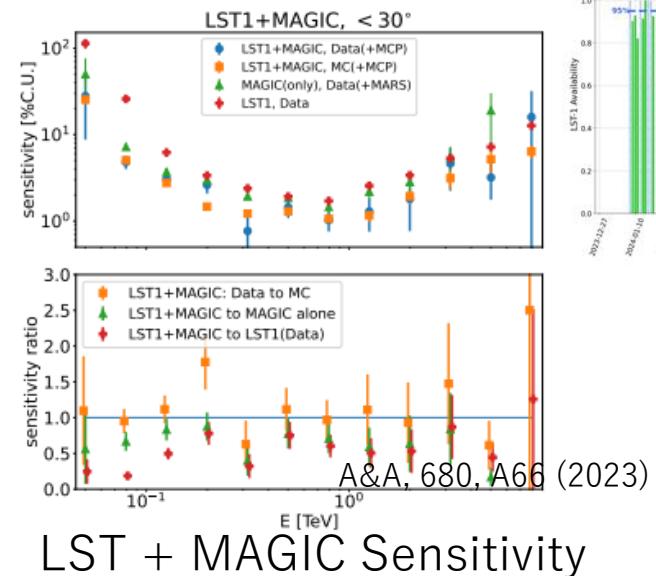
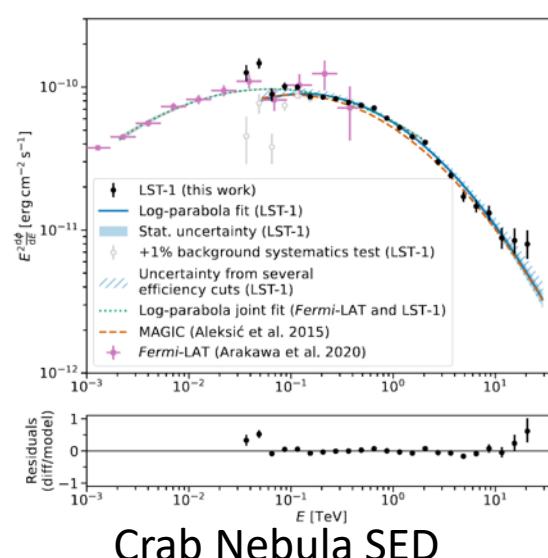
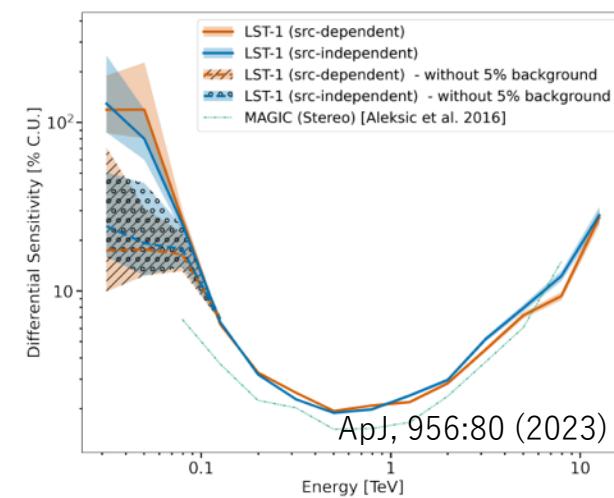
Mirror@LaPalma

	Arch	Mirror	Camera
LST1		operational	
LST2	25/6	25/8	26/03
LST3	24/1 done	25/5	25/9
LST4	24/8 done	24/12 done	25/5

4-LST array obs. starts in 2026



- More than 2700 hours since 2020
 - 30-40% of the time with MAGIC-Joint
 - 90% availability



Availability

LST-1 Sensitivity

Crab Nebula SED

LST + MAGIC Sensitivity

Total
125 # total alerts
20 # observed alerts
46.7h observed

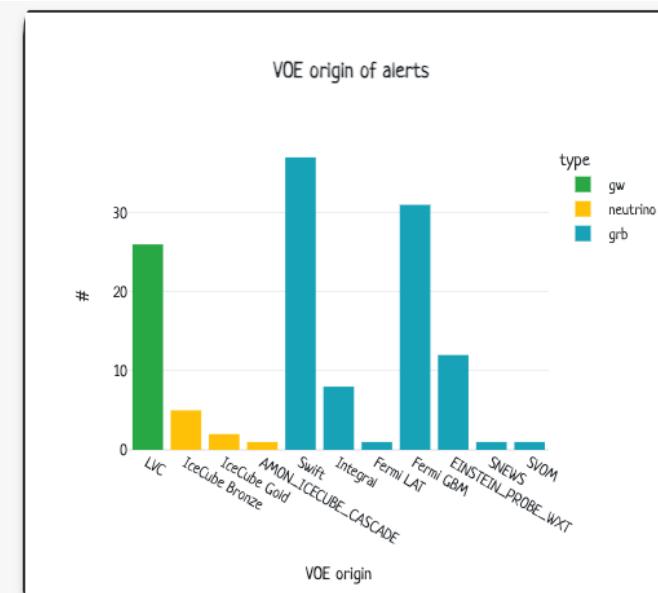
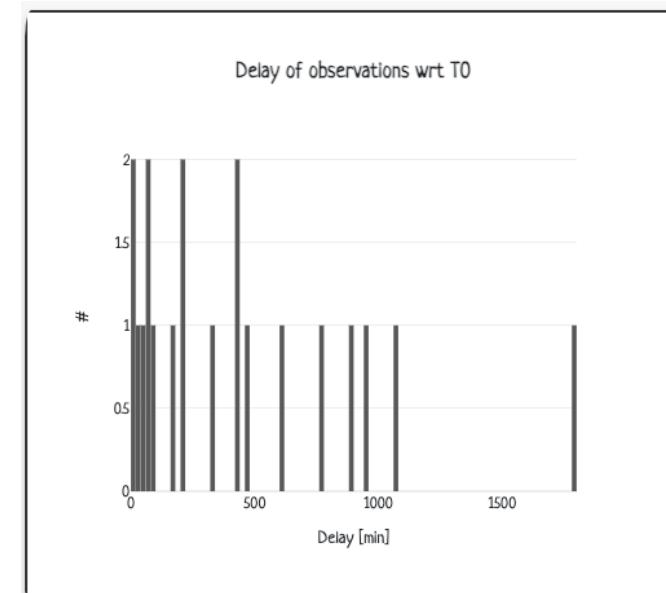
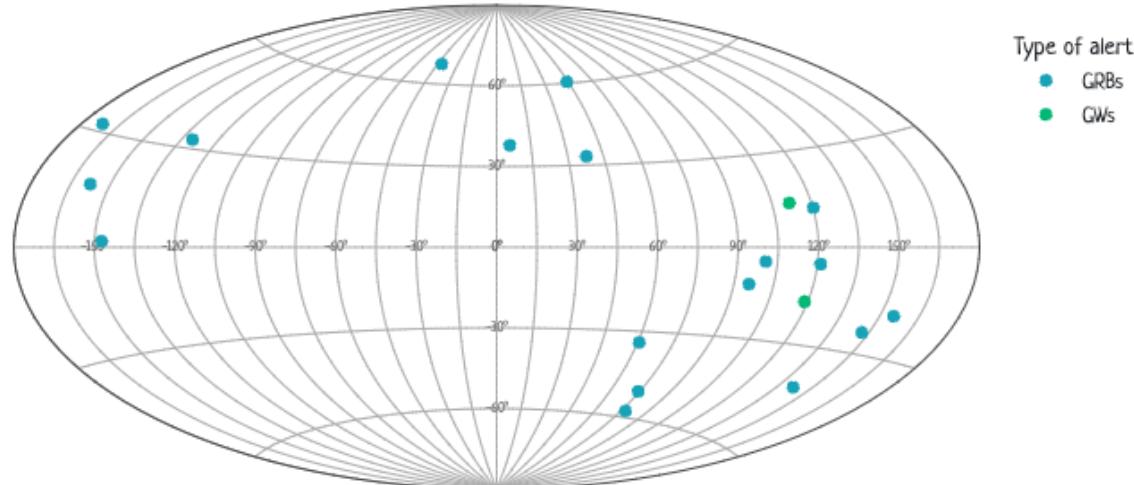
GRBs
91 # total alerts
18 # observed alerts
43.4h observed

GWs
26 # total alerts
2 # observed alerts
3.3h observed

Neutrinos
8 # total alerts
0 # observed alerts
0.0h observed

- Fast rotation (180deg/20sec)
- Automatic Reaction on alert
- No detection so far

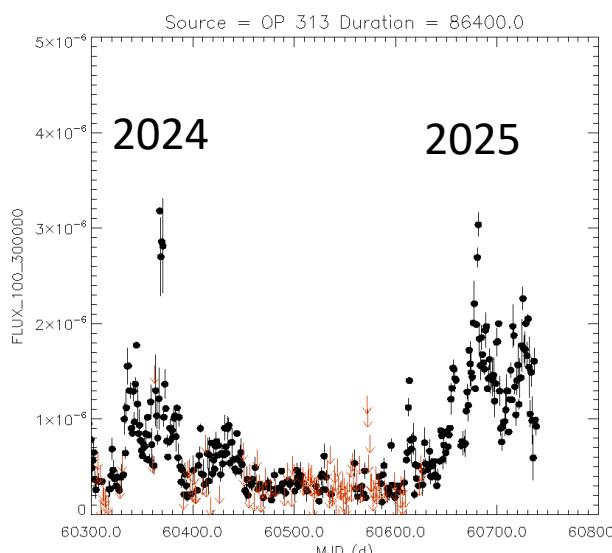
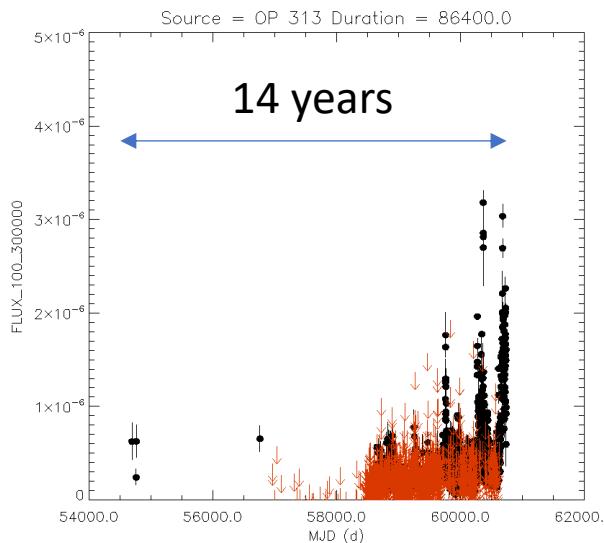
Sky map in Galactic coordinates of all observed alerts



- Brightest of All Time GRB. $Z = \sim 0.151$, $E^{\text{iso}} \sim 10^{55}$ erg

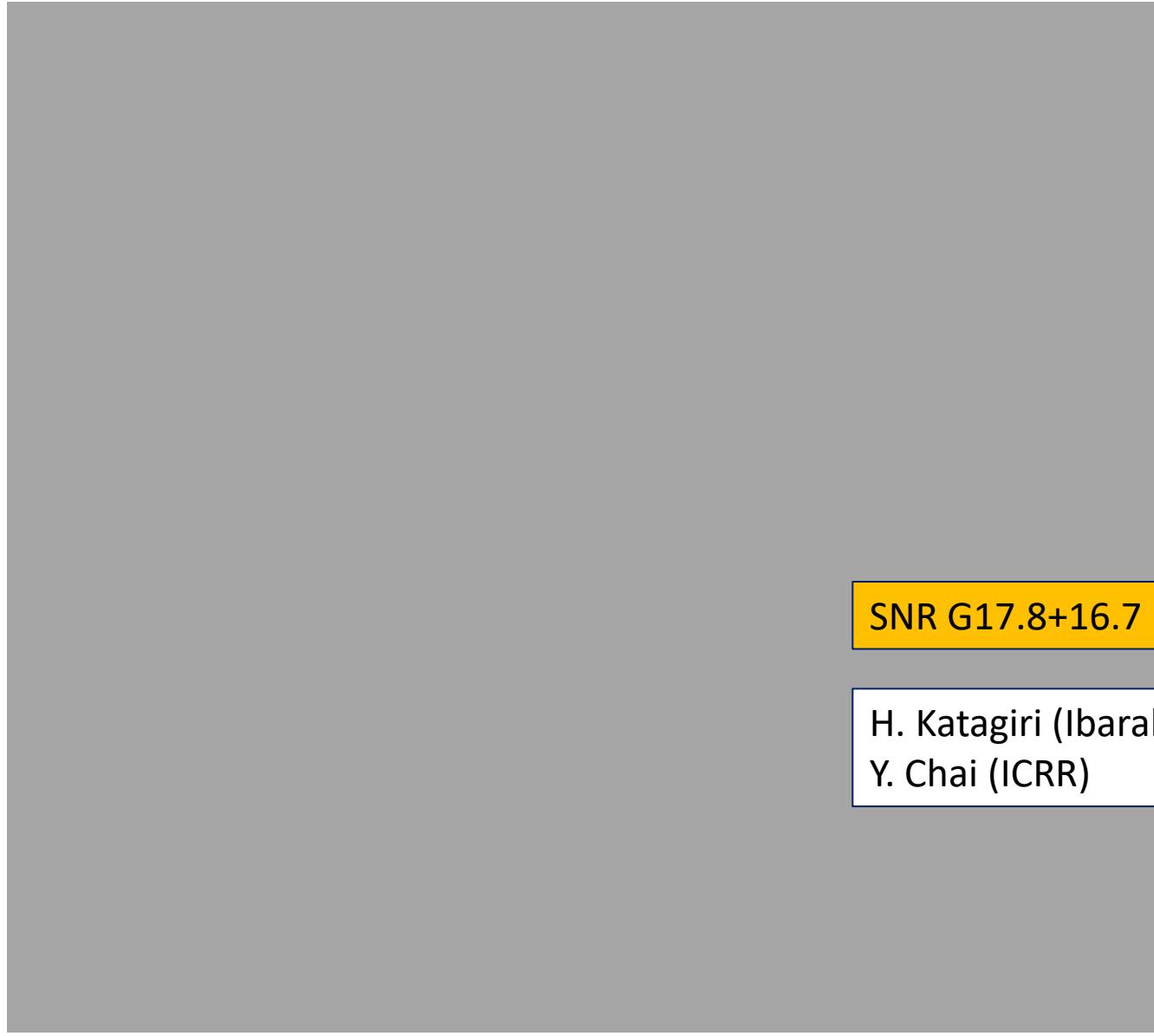
- Observation started at T0+1.3d under very strong moon
- To bypass the “Safety Limit”, PMT HVs had to be reduced. Analysis was very difficult.
- 4 sigma excess could be obtained
- Consistent with LHAASO extrapolation after break.

- **Very distant FSRQ** $Z = 0.997$, the farthest AGN detected with IACT



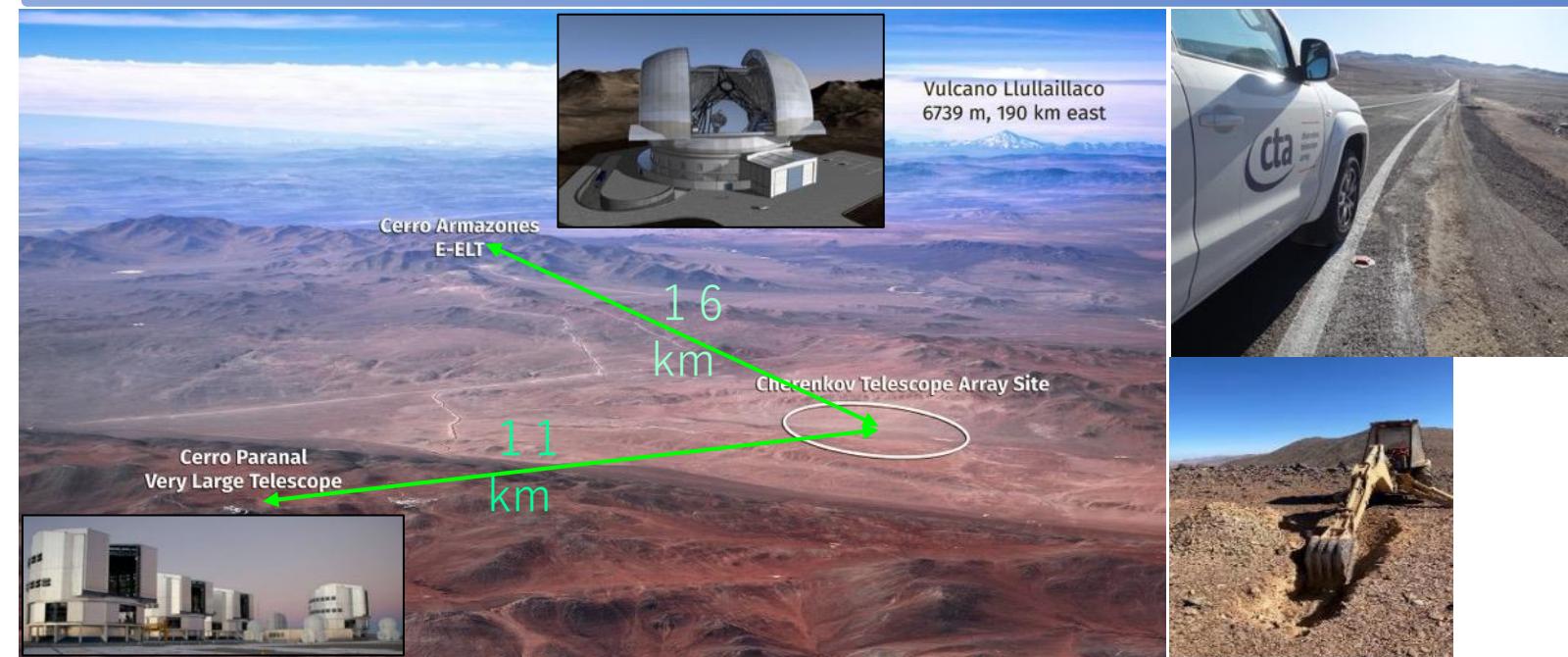
J. Baxter (ICRR)
S. Nozaki (ICRR, MPI)

- Very important for EBL study
- Very active again inand keep observing

**SNR G17.8+16.7**

H. Katagiri (Ibaraki)
Y. Chai (ICRR)

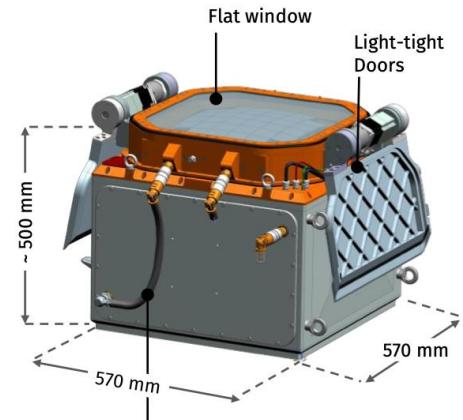
**I. Vovk (ICRR)****Cygnus Cocoon Scan**



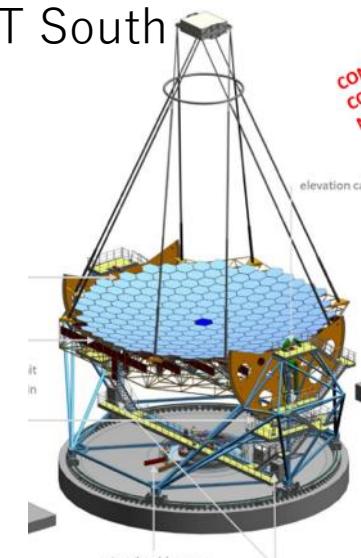
MST camera on HESS



SST Camera



LST South



□ Infrastructure

- Array road construction till July 2026
- Power for first tels by Feb 2026
- Tentative data center by Feb 2026
- Foundation of 5 SST and 2 MST by Jan 2026

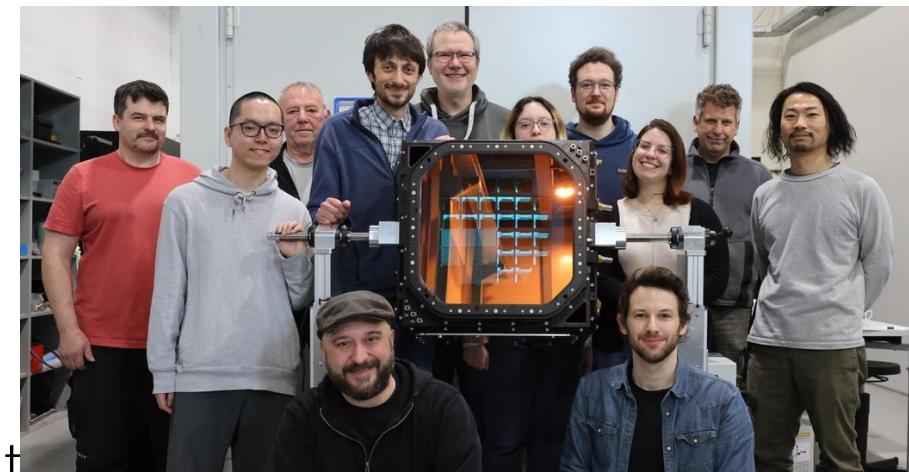
□ Telescope elements

- LST
 - Copy of North Camera
 - New Structure designed.
 - New mirror developed.
- MST
 - Prototype tested.
 - Camera mounted on HESS and validated.
- SST
 - Prototype Telescope validated.
 - Prototype Array being constructed in Tenerife.

5 SST, 2 MST, 2 LST will be ready in 2027 11

❖ **Small-sized telescope started production of the first telescope**

- Nagoya group is in charge of SiPM for SST camera.
 - 224 SiPM modules (14,336 channels, worth 5 SSTs) are being procured by Nagoya University
 - Currently, commissioning a quadrant of the camera



❖ **Dark current of 64 channels of SiPMs is measured for more than 3 months**

- One SiPM suffered $\times 10$ dark current increase after 3 months
 - The dark count rate is similar to typical SiPM, which indicates large dark current does not originate from SiPM APDs
 - Visual inspection and temperature measurements found this is caused by a damage to the electrode. Automatic visual inspection will be added to screen out this kind of defects in the future.

- CTAO is next generation gamma-ray observatory, in both hemispheres.
 - Cosmic Ray origin, BH/NS physics, Cosmology/Basic physics
 - **CTAO ERIC** is established on Jan 2025, which will accelerate the construction.
 - Both arrays will be completed by 2028
- North Site construction is going well.
 - LST array will be **completed in this year**
- LST-1 is producing lots of scientific results
 - Recent ones include BOAT GRB, OP313, SNR G17.8 etc
 - Automatic Transient follow-up observation is on going
- South Site is also progressing well.
 - Starting from Infrastructure.
 - 5 SST, 2 MST, 2 LST will be ready in 2027